

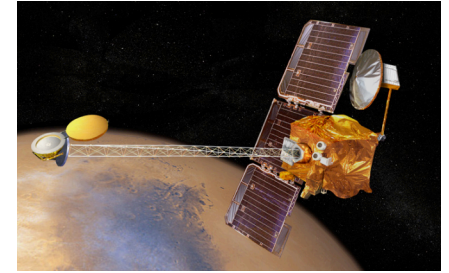
Comments about MSL Landing sites:

***Samara Vallis, Eos Chasma, Aeolis Region and
Shalbatana Vallis regions***

Igor Mitrofanov for DAN/MSL Science Team

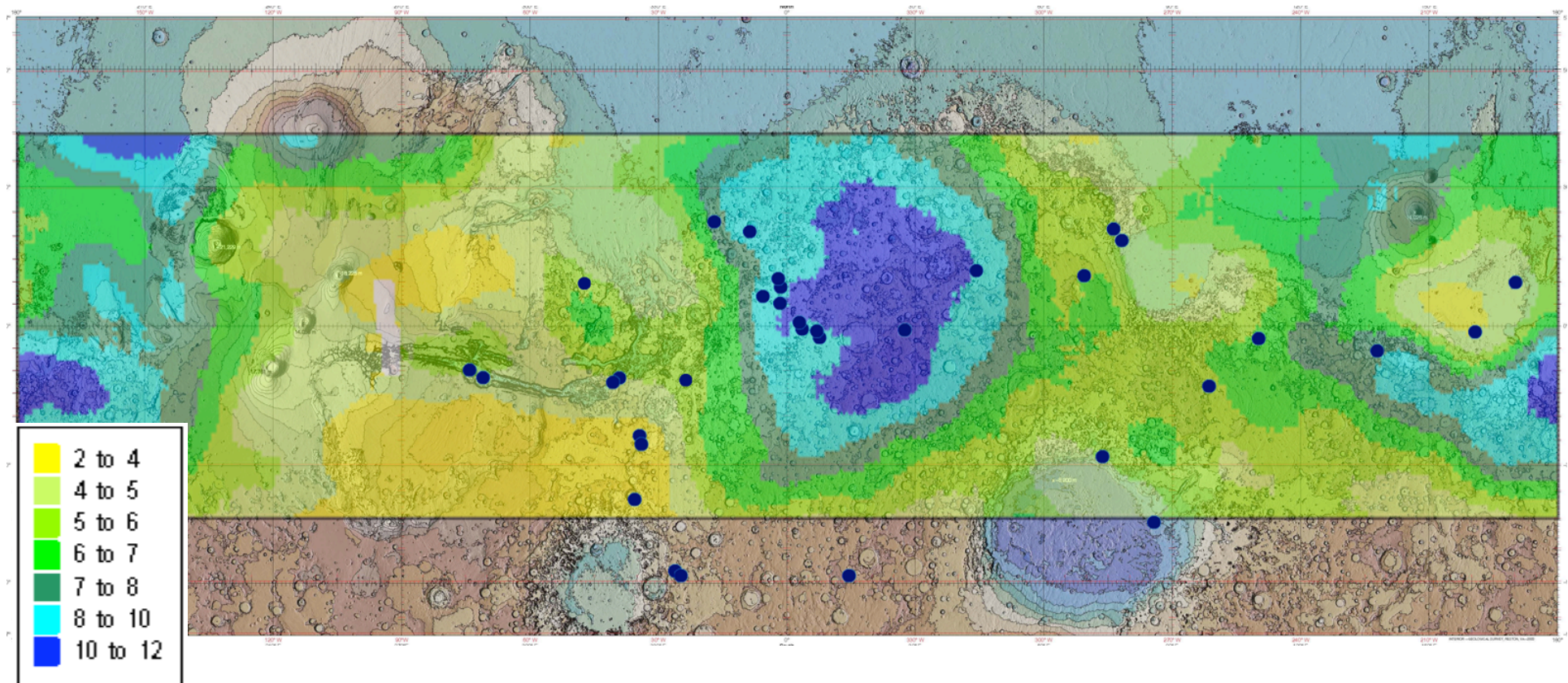
***MSL Landing Site selection Workshop #2
Pasadena, October 23-25, 2007***

Introductory Remarks: Following the Water

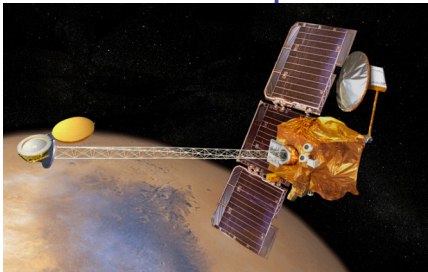


Map of the HEND water equivalent (%) based on 2-layer soil model with 2% of water amount in upper 5-cm layer.

● - landing sites proposed on the first MSL Workshop



MSL LS	HEND water equivalent (%)													
	1	2	3	4	5	6	7	8	9	10	11	12	13	14
1. E. Meridiani														
2. Meridiani Crater														
3. Becquerel Crater														
4. SW Arabia Terra														
5. W. Arabia														
6. N Meridiani														
7. W. Meridiani														
8. Meridiani Bench														
9. Marwth Vallis														
10. Gale Crater														
11. Iani Chaos														
12. W Candor East														
13. Melas Chasma														
14. NE Syrtis Major														
15. Margaritifer basin														
16. Elysium/Avernus Colles														
17. Nili Fossae Crater														
18. Terby Crater														
19. Xanthe/HypanisVallis														
20. Hellas/Dao Vallis														
21.E. Melas Chasma														
22. Eos Chasma														
23. Athabasca Vallis														
24. Nilo Syrtis														
25. Juventae Chasma														
26. Isidis Basin Escarp														
27. SML Crater														
28. NW Slope Valleys														
29. Nili Fossae Trough														
30. Holden Crater Fan														
31. Eberswalde Crater														
32. Gullies/Hale Crater														
33. Argyre														



Introductory Remarks: How we consider our new sites -

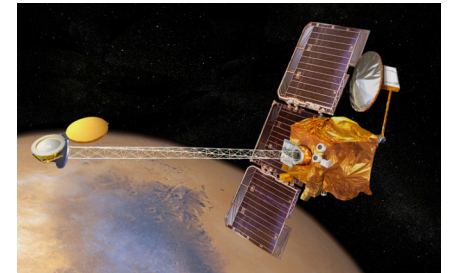
- (1) We want to follow the water, as the most promising sites of habitability**
- (2) We want to correlate the *evidence* for water in present with the *signatures of water* in the past**
- (3) We want to continue studies of our candidates using all new observational data and new concepts and paradigms**
- (4) We will support our studies by engineering analysis**

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Samara Vallis, Eos Chasma, Aeolis Region, and Shalbatana Vallis Regions

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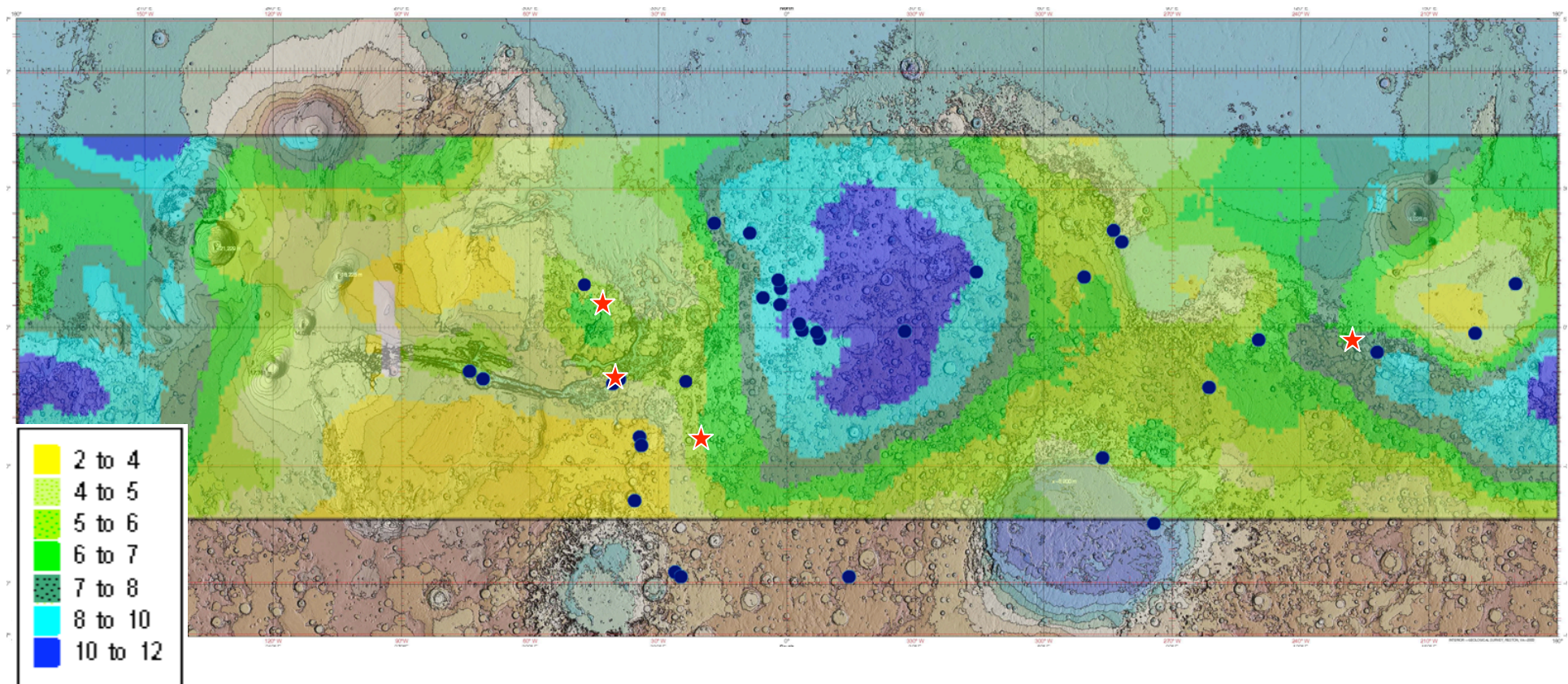
Introductory Remarks: Following the Water



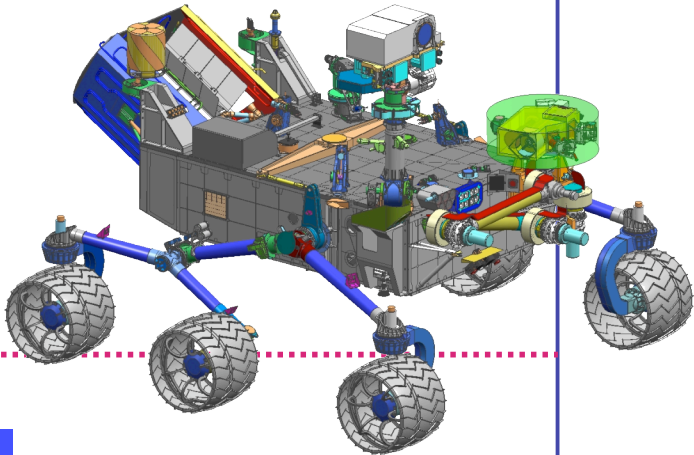
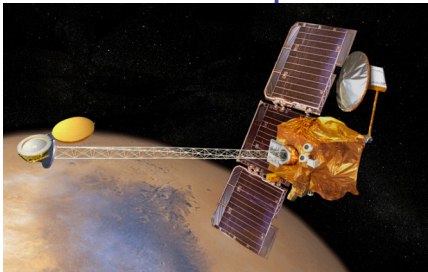
Map of the HEND water equivalent (%) based on 2-layer soil model with 2% of water amount in upper 5-cm layer.

● - landing sites proposed on the first MSL Workshop

★ - landing sites proposed in this talk



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33. Argyre														
34. Samara Vallis Basin														
35. Eos Chasma Alluvial														
36. Aeolis fan-delta														
37. Shalbatana V. region														

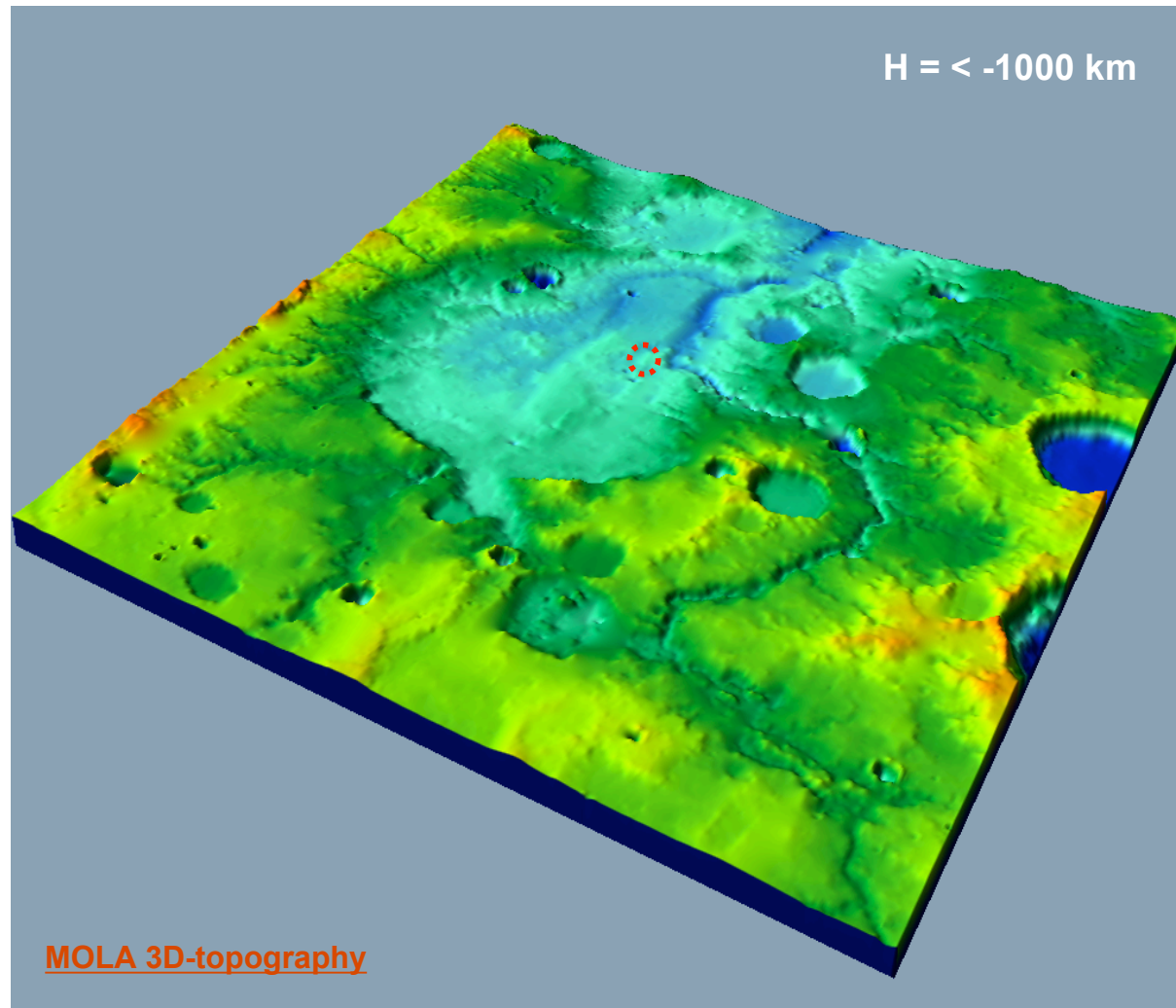


Site from Ruslan Kuzmin and DAN Science Team:

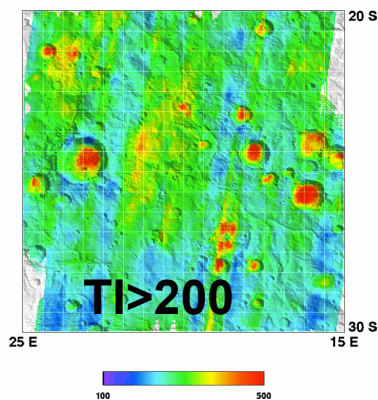
Samara Vallis

(as Fluvio-Lacustrine Basin)

Crater Moroz-Samara Vallis (as Fluvio-Lacustrine Basin)



TES TI, Ls 100-180



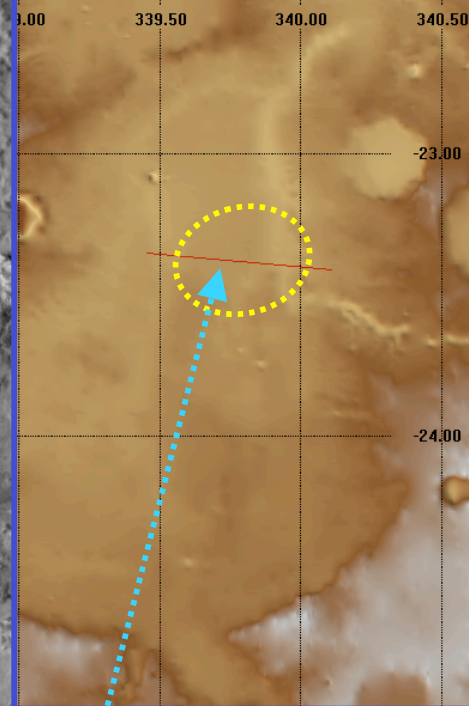
Samara Vallis basin

Age: Noachian

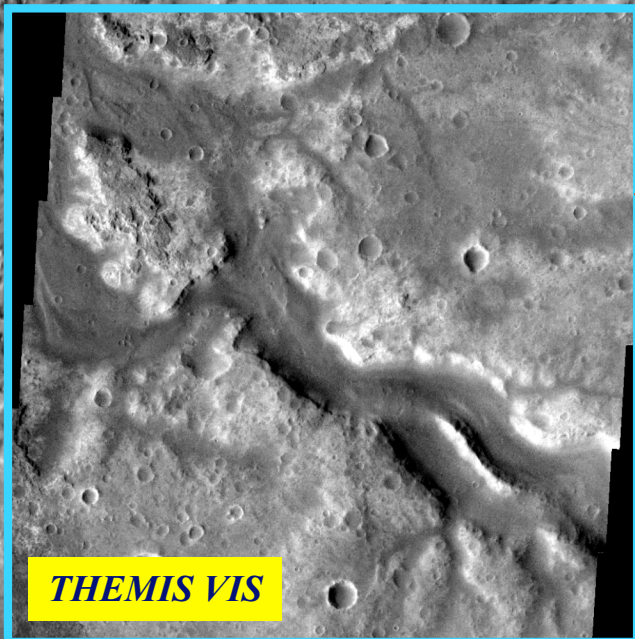
THEMIS Td mosaic

Samara Vallis

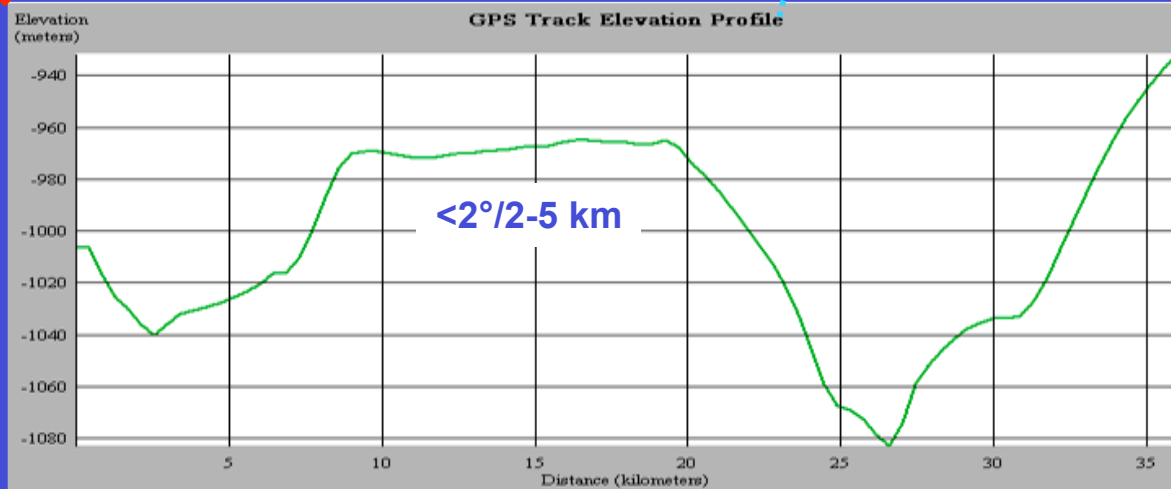
LS

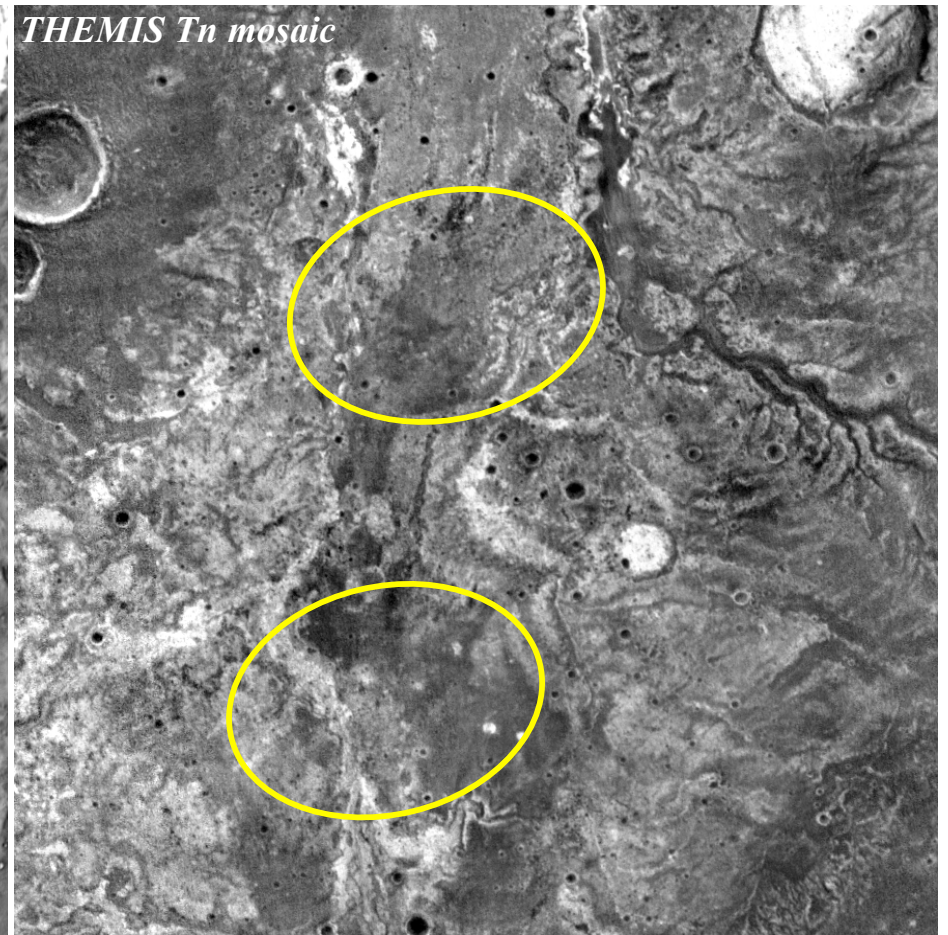
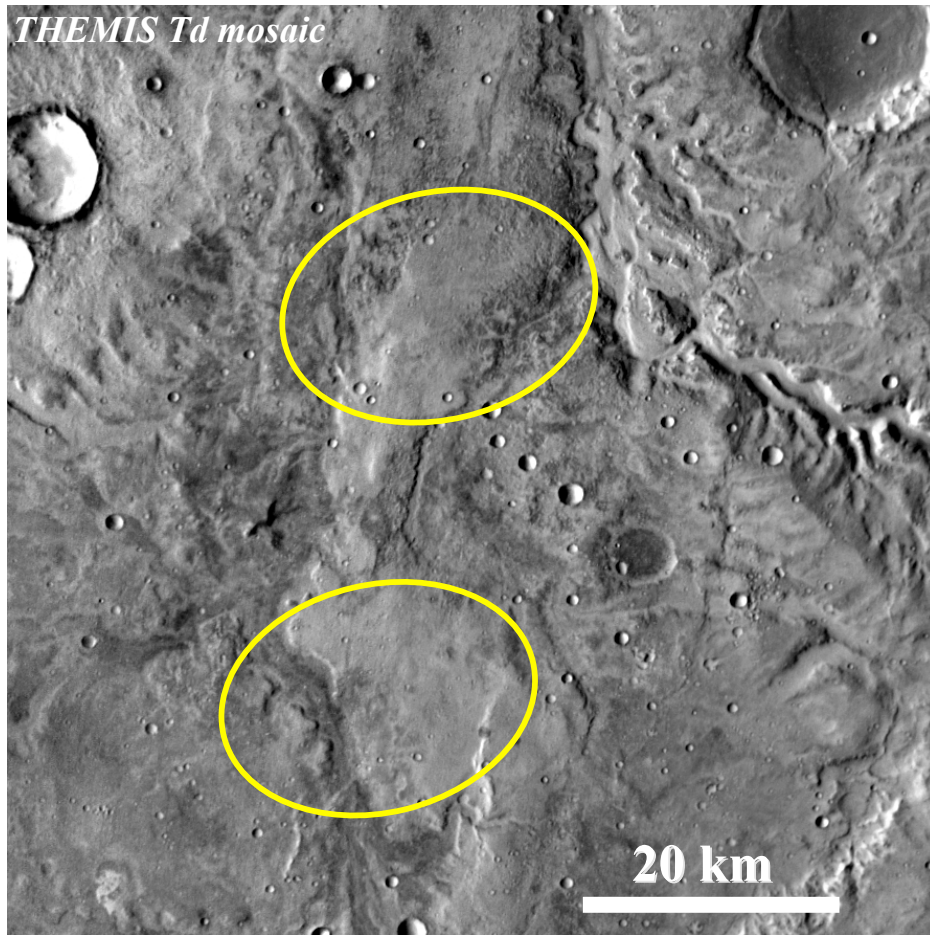


MOLA, 128 px /deg



THEMIS VIS





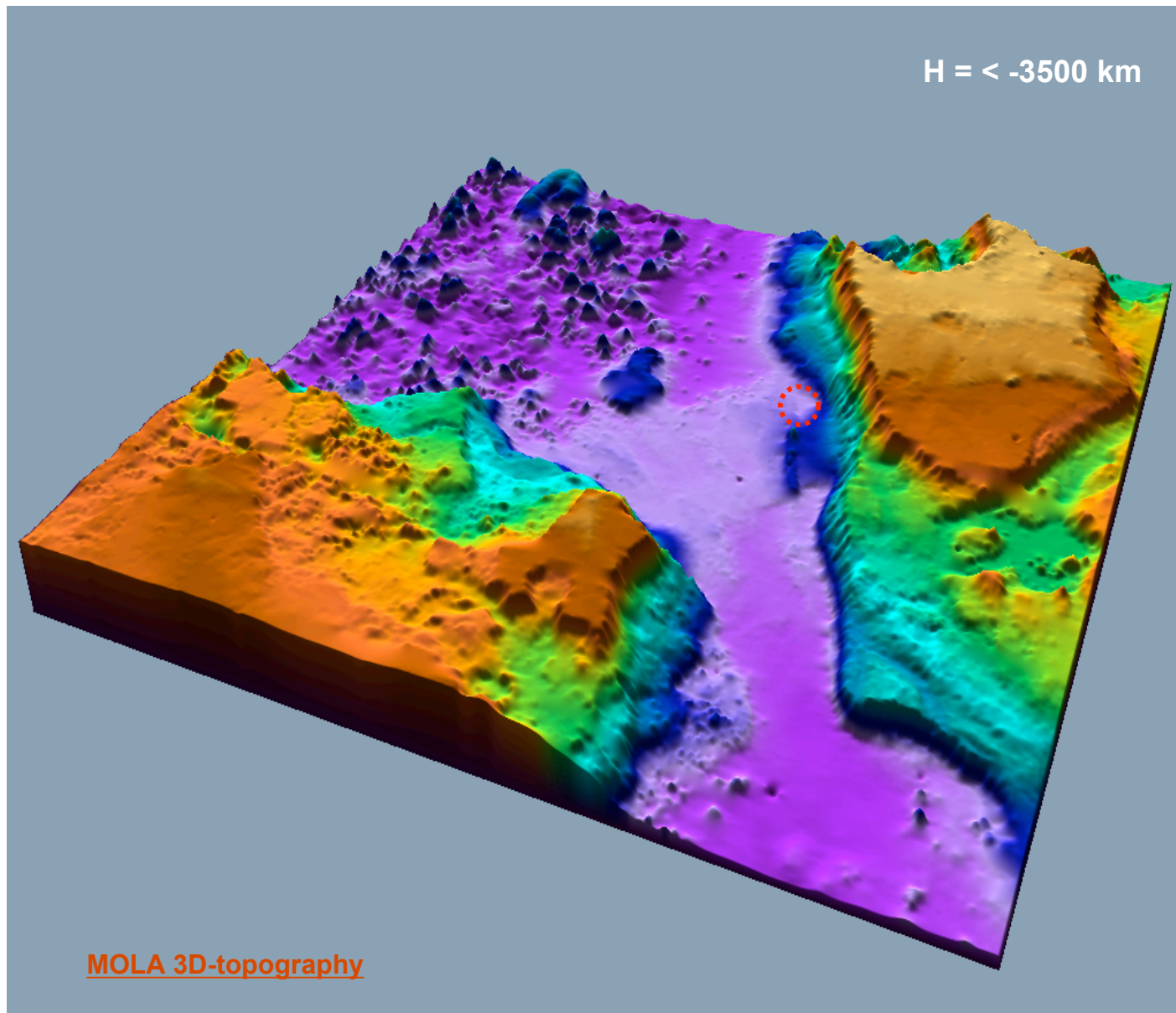
Comparison of the THEMIS Tday and Tnight images show that two potential MSL LS locations within Samara Vallis's Basin are characterized by diversity of a fluvio-lacustrine lithological facies.

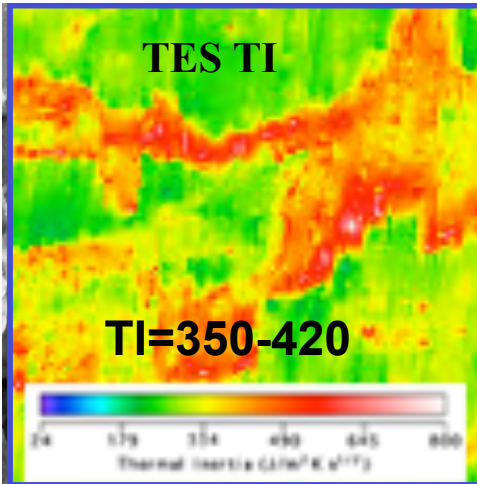
Site from Ruslan Kuzmin and DAN Science Team:

Eos Chasma

(as Alluvial Fans)

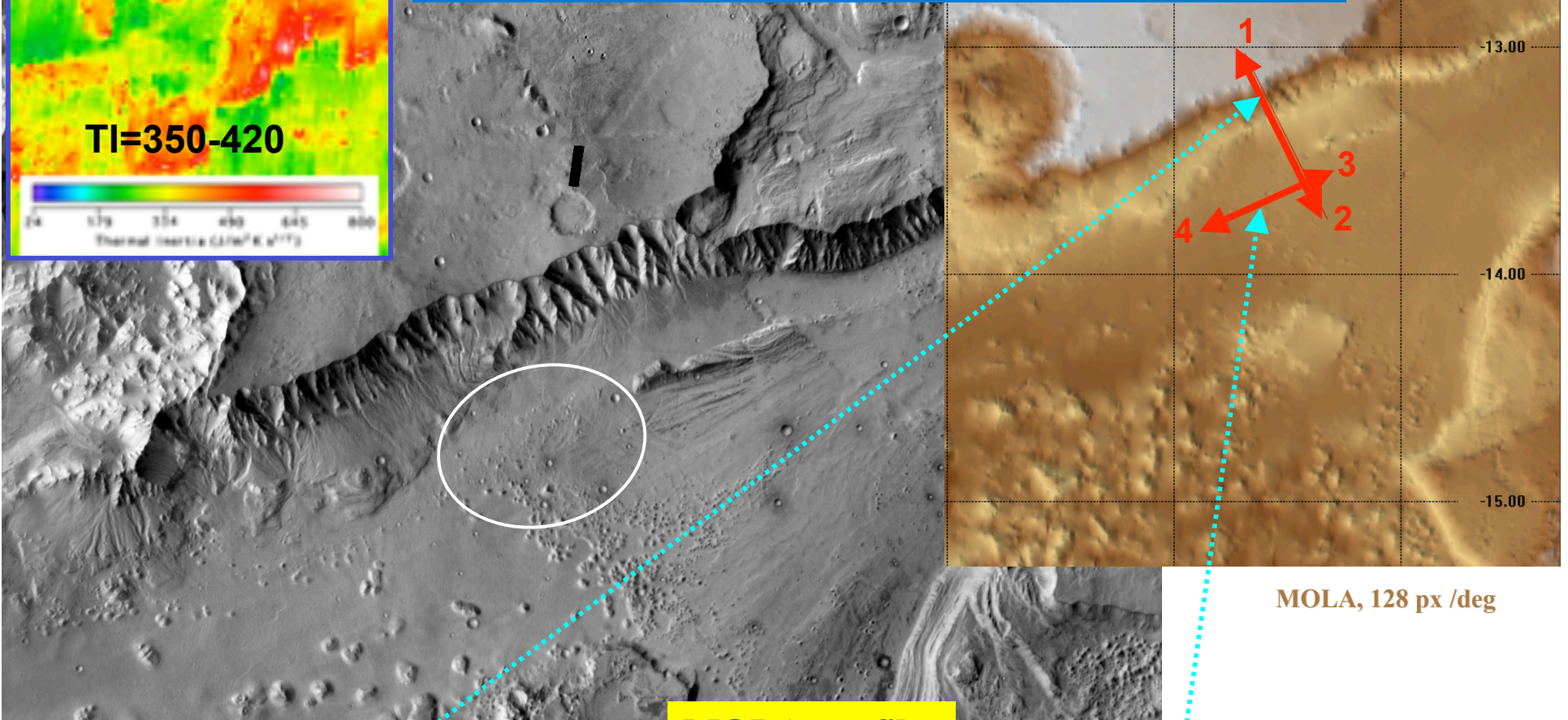
Eos Chasma (Alluvial Fans)



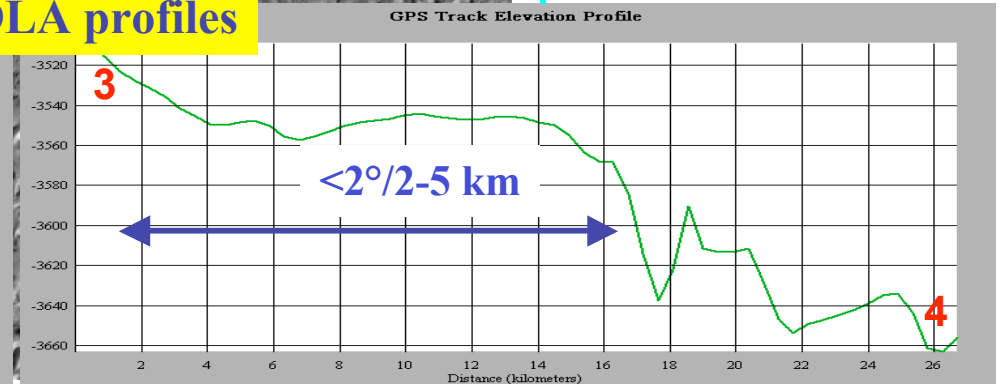
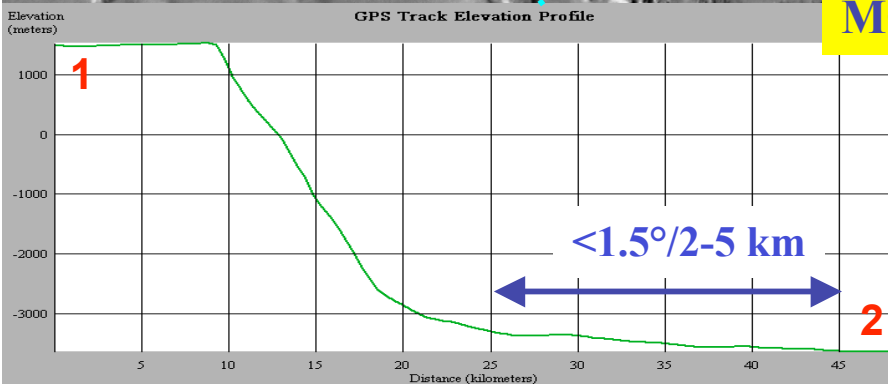


Alluvial fans in Eos Chasma

Age: Late Hesperian-early Amazonian

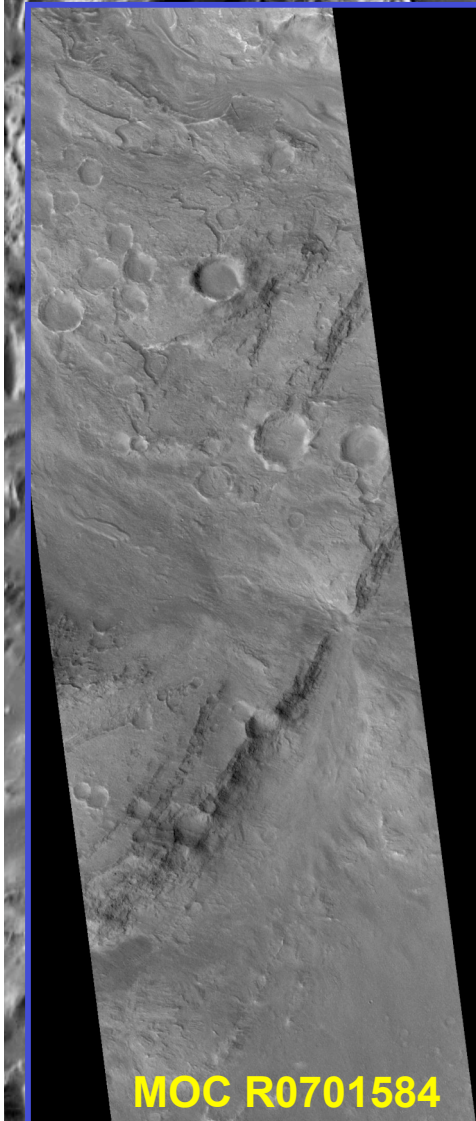


MOLA profiles

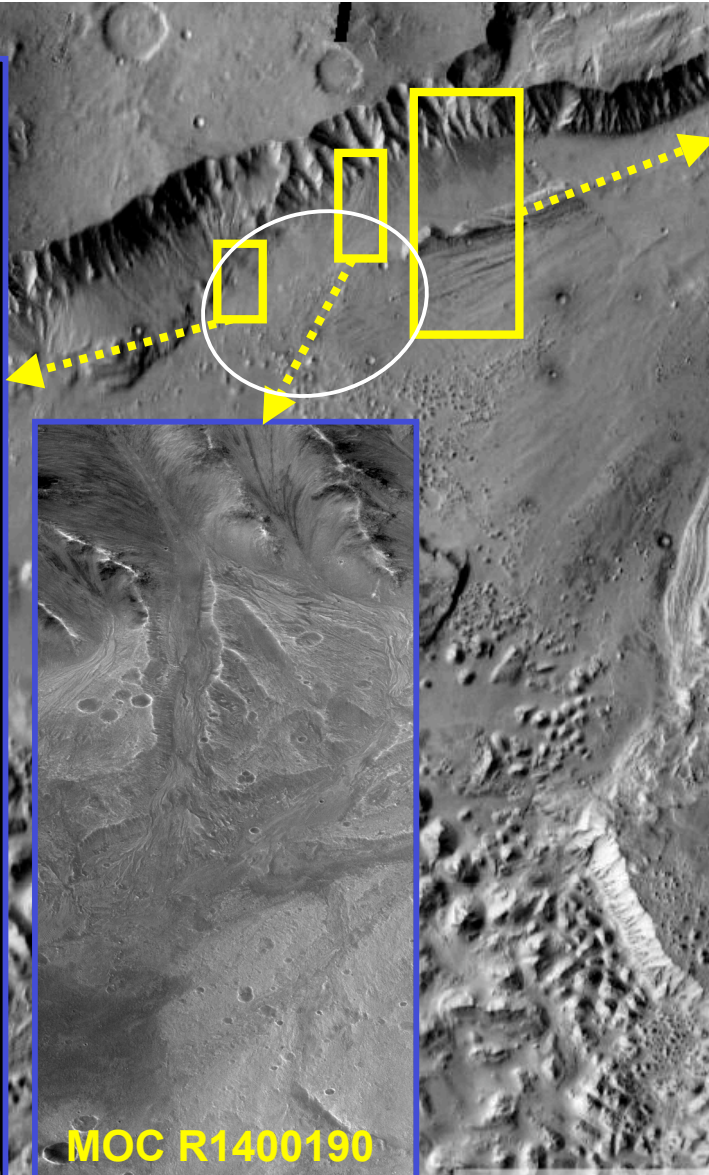


THEMIS mosaic

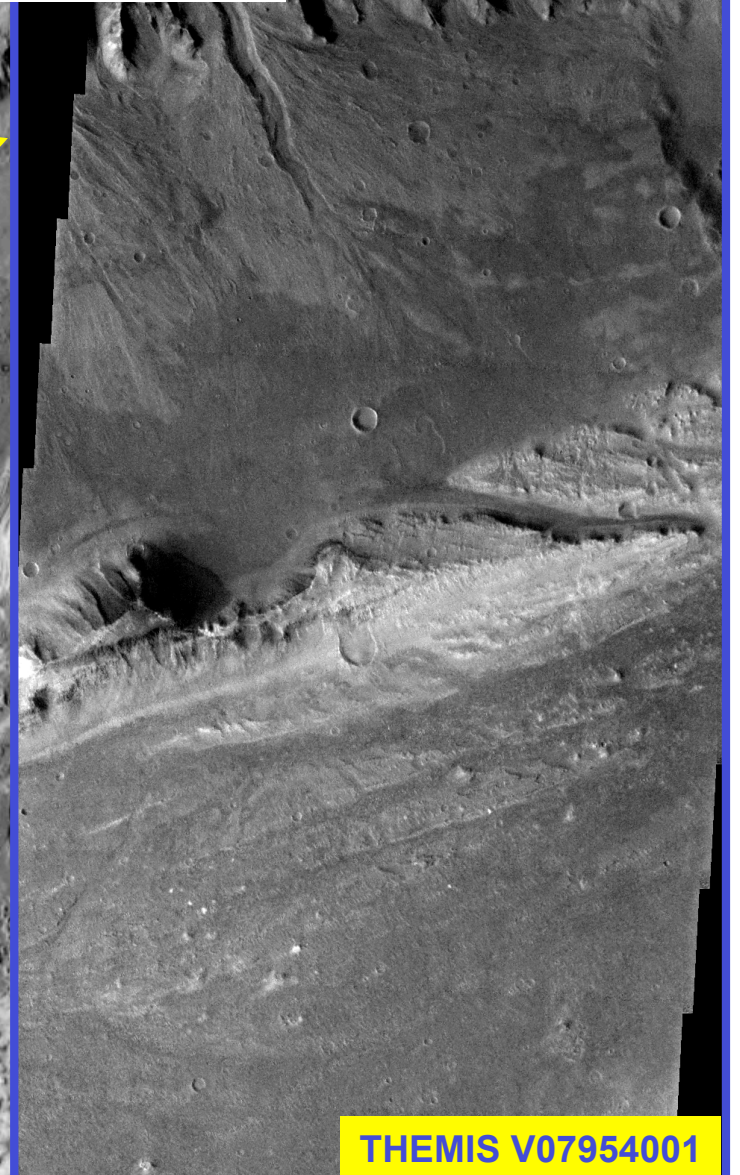
**Great number of the alluvial fans
and stream's erosion features within
LS ellipse.**



MOC R0701584



MOC R1400190



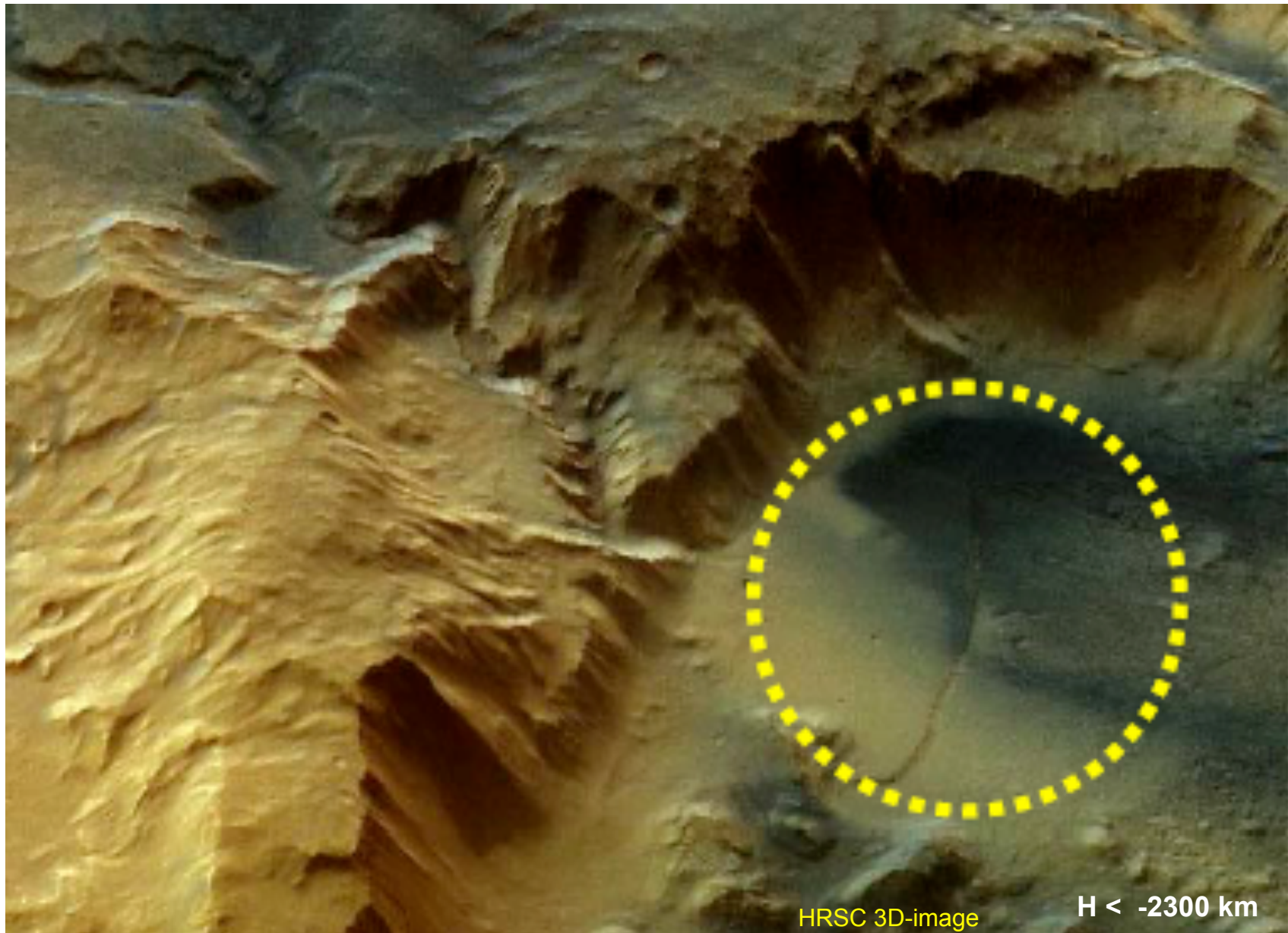
THEMIS V07954001

Site from Ruslan Kuzmin and DAN Science Team:

Aeolis Region

(as Lobate Fan-Delta)

Aeolis region

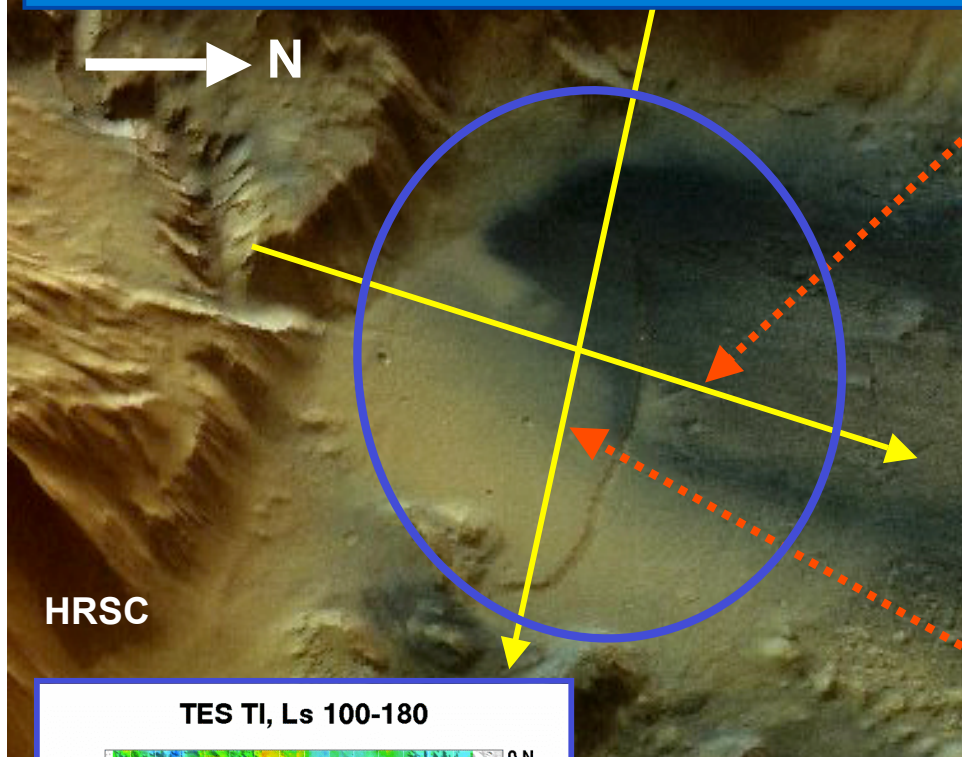


HRSC 3D-image

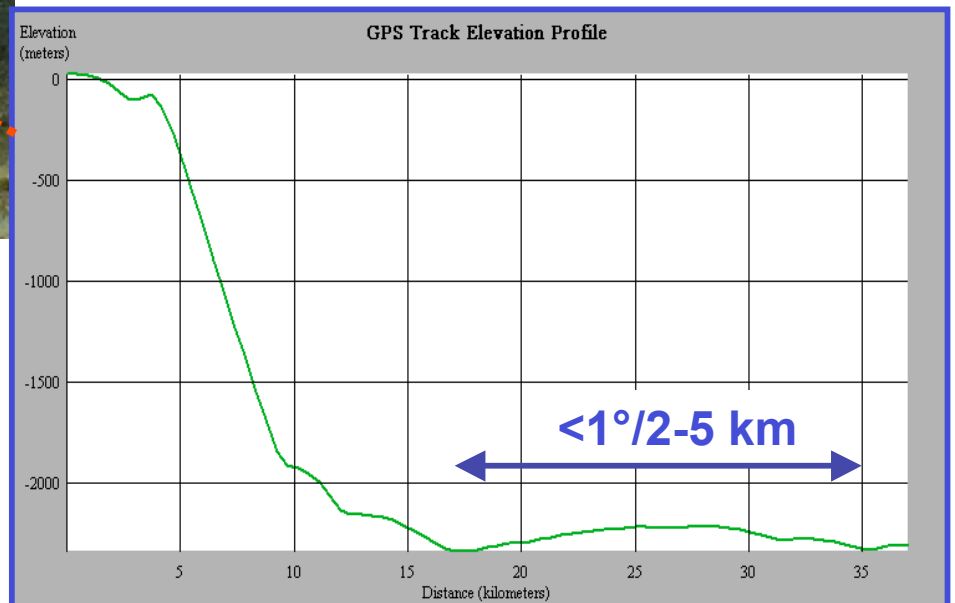
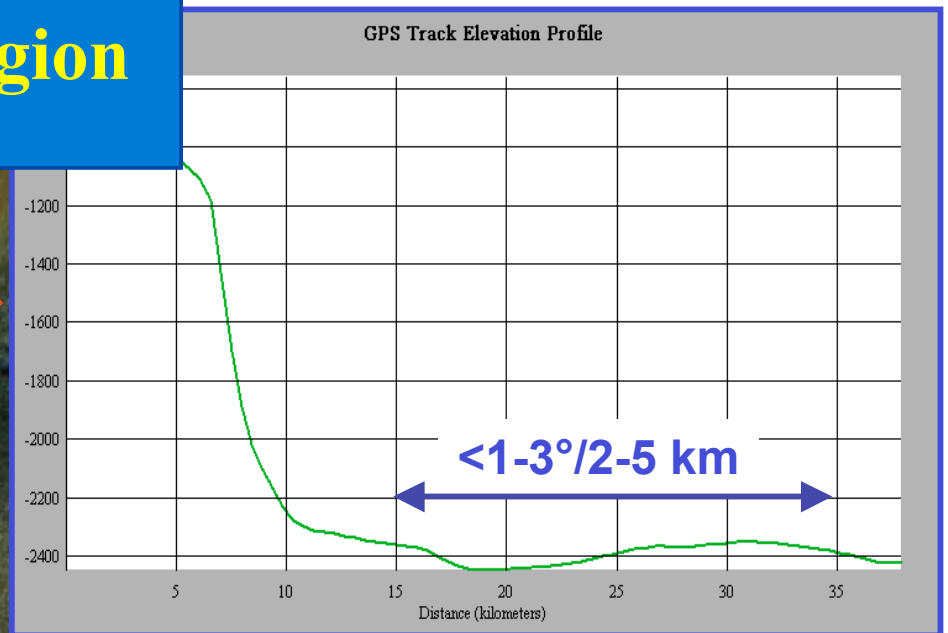
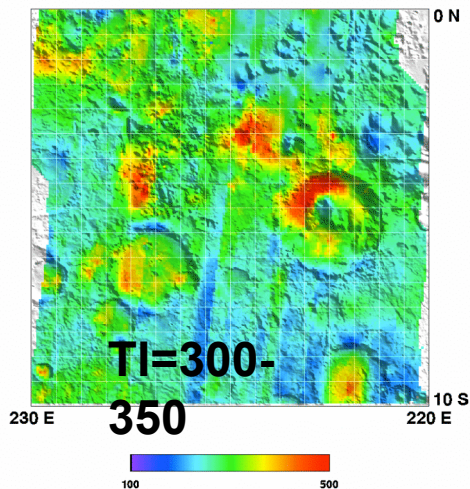
H < -2300 km

Lobate fan-delta in Aeolis region

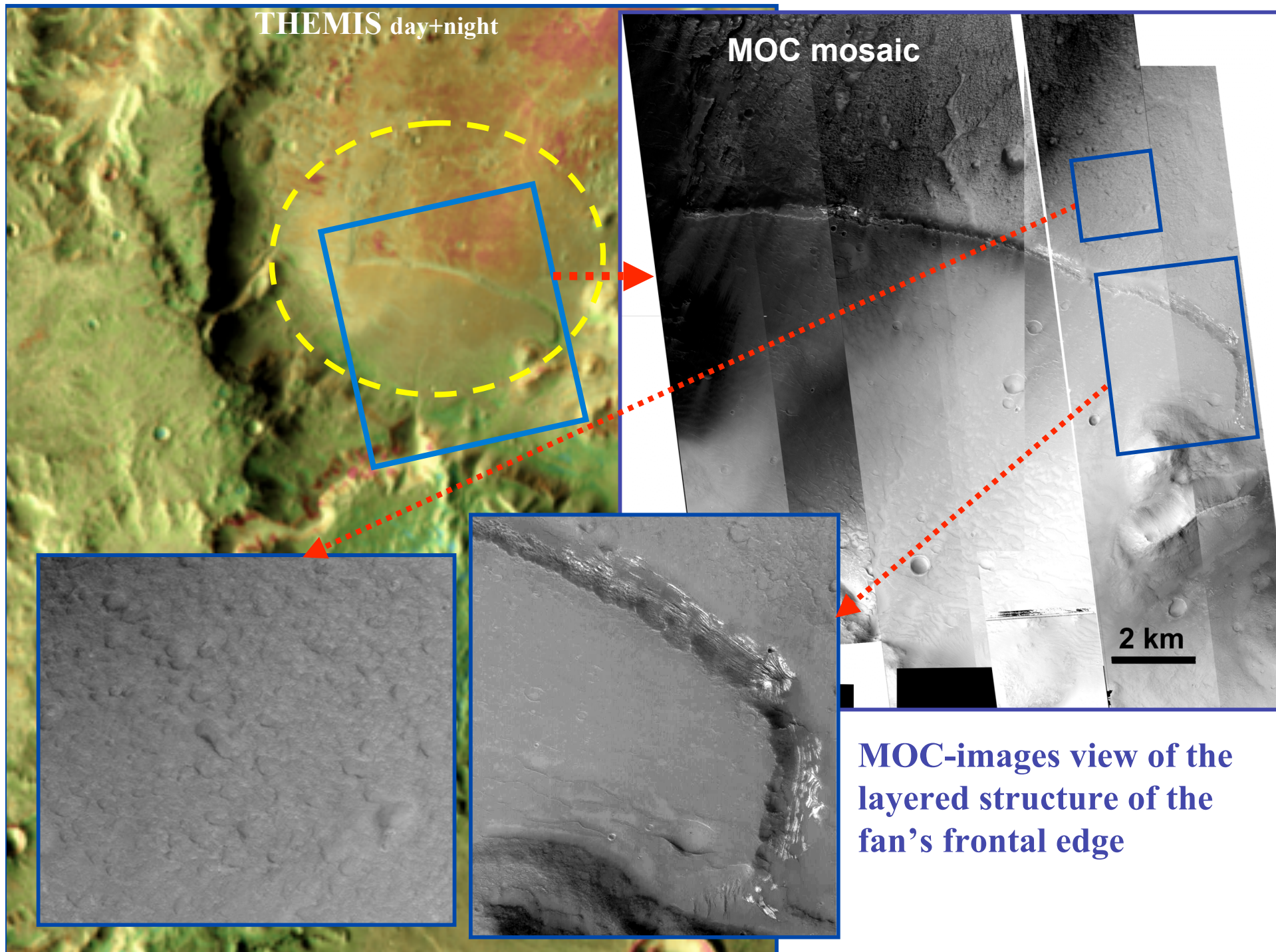
Age: Noachian-Early Hesperian

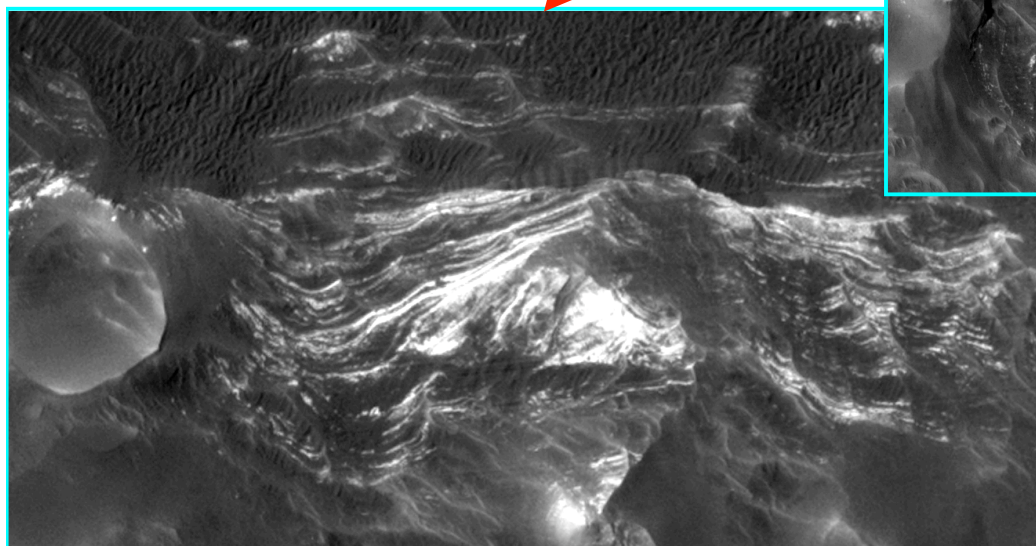
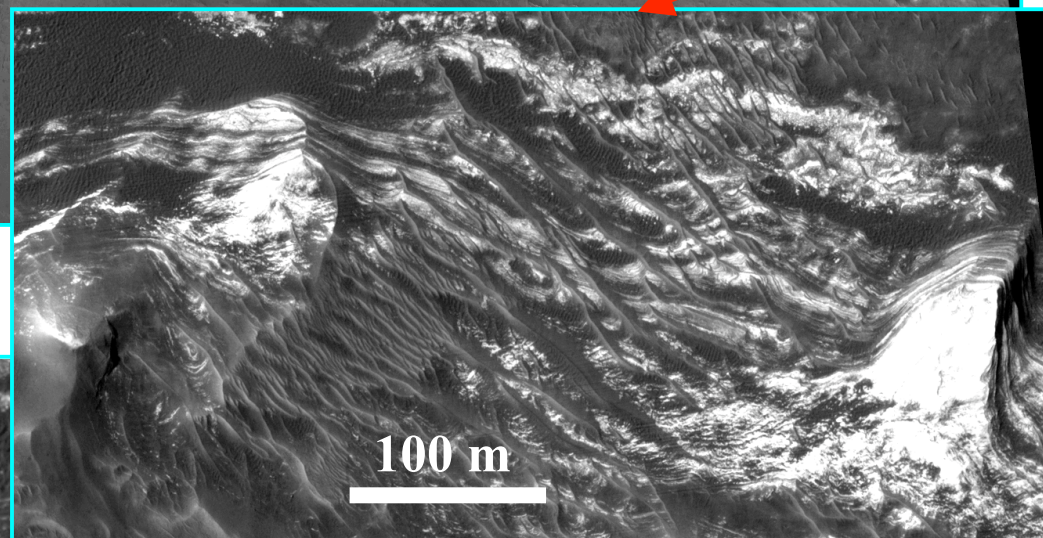
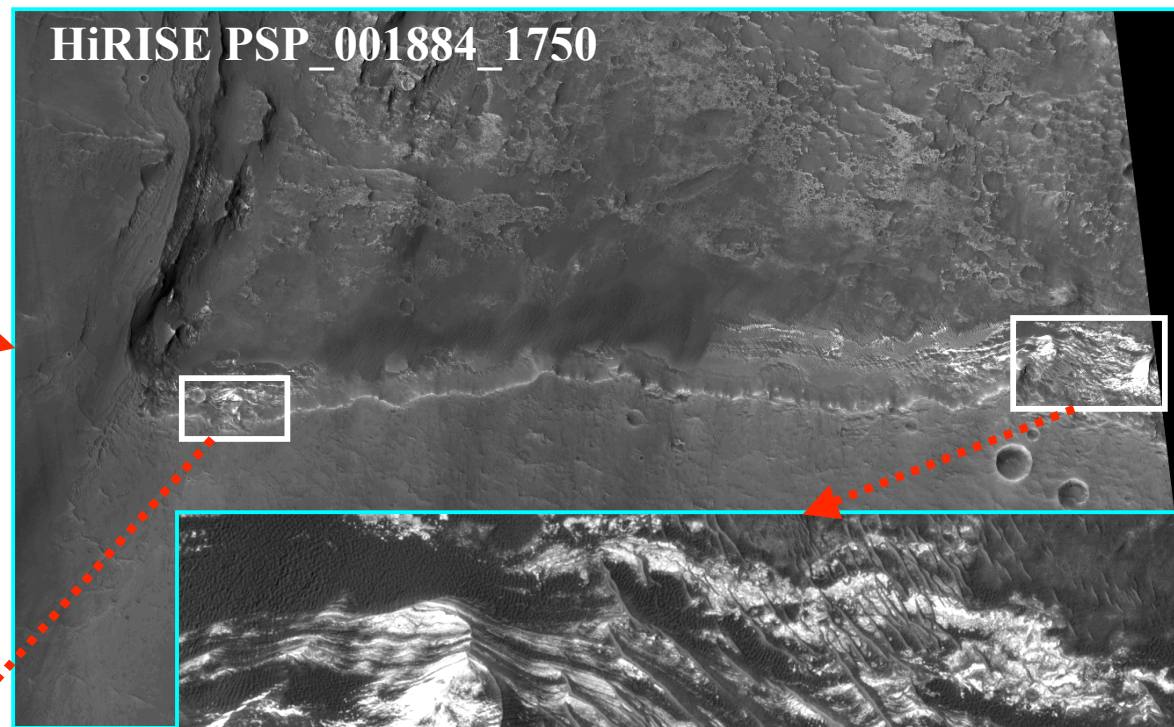
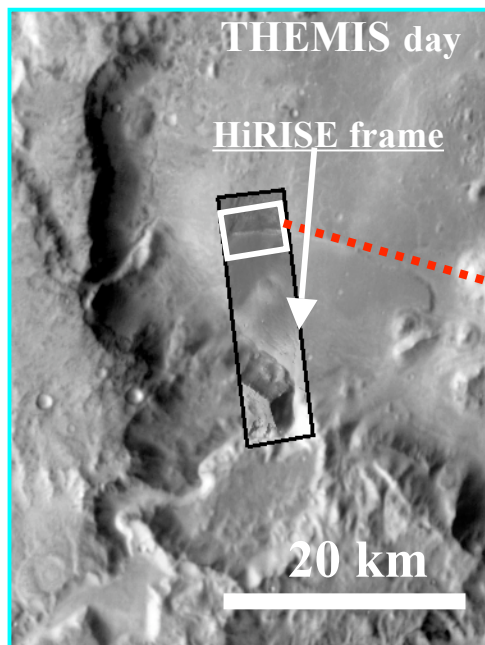


TES TI, Ls 100-180



MOLA profiles along and across of the fan





Outcrops of the fine layered fan-delta deposits on the fan's frontal edge and underlying sediments

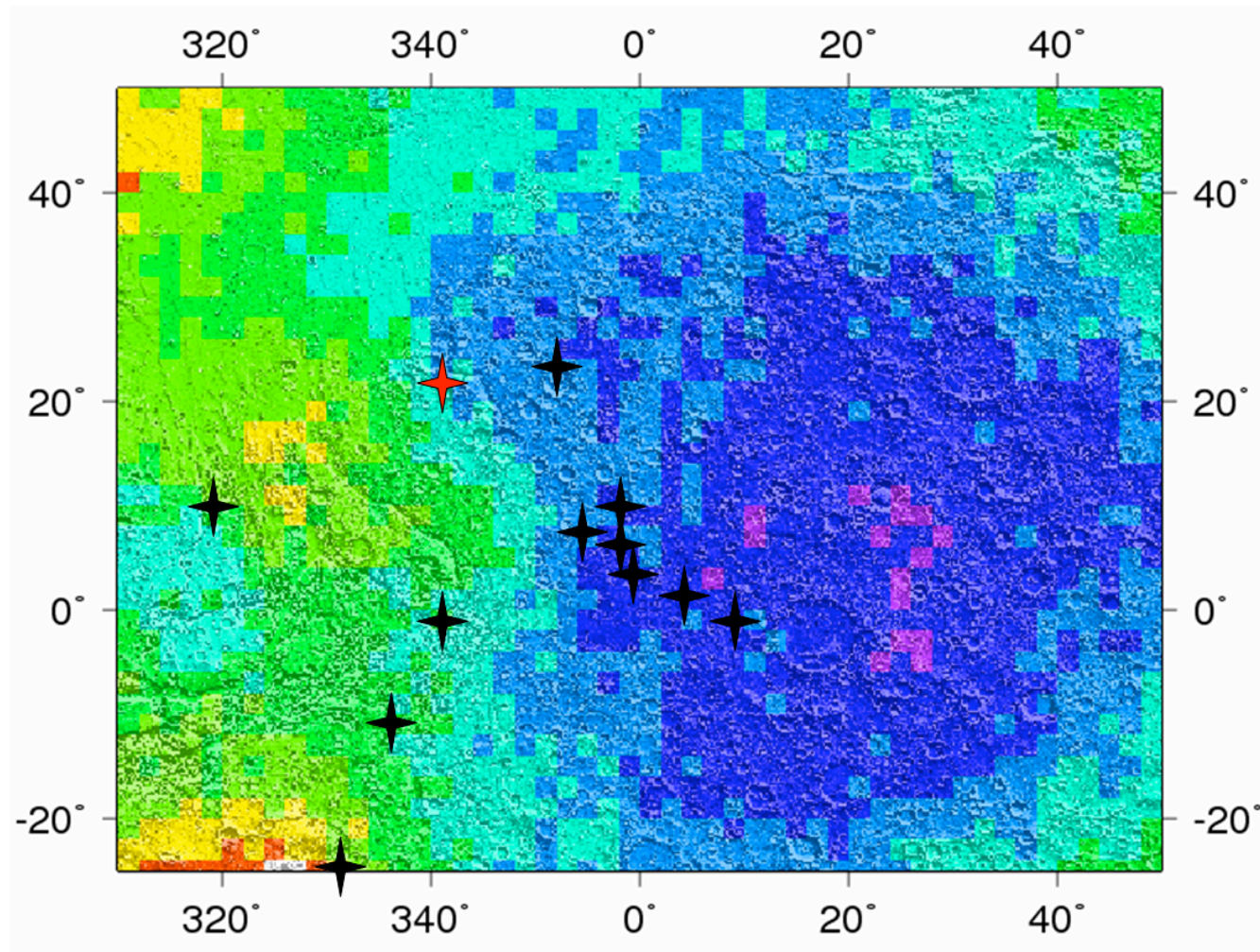
**Site from Nikita Demidov, Alberto Behar, Igor
Mitrofanov and DAN Science Team:**

Shalbatana Vallis Regions

(as current water site + Phillosilicate site)

Arrabia Terra - is it region for current water site + Phillosilicate site?

Arabia Terra HEND neutron flux map

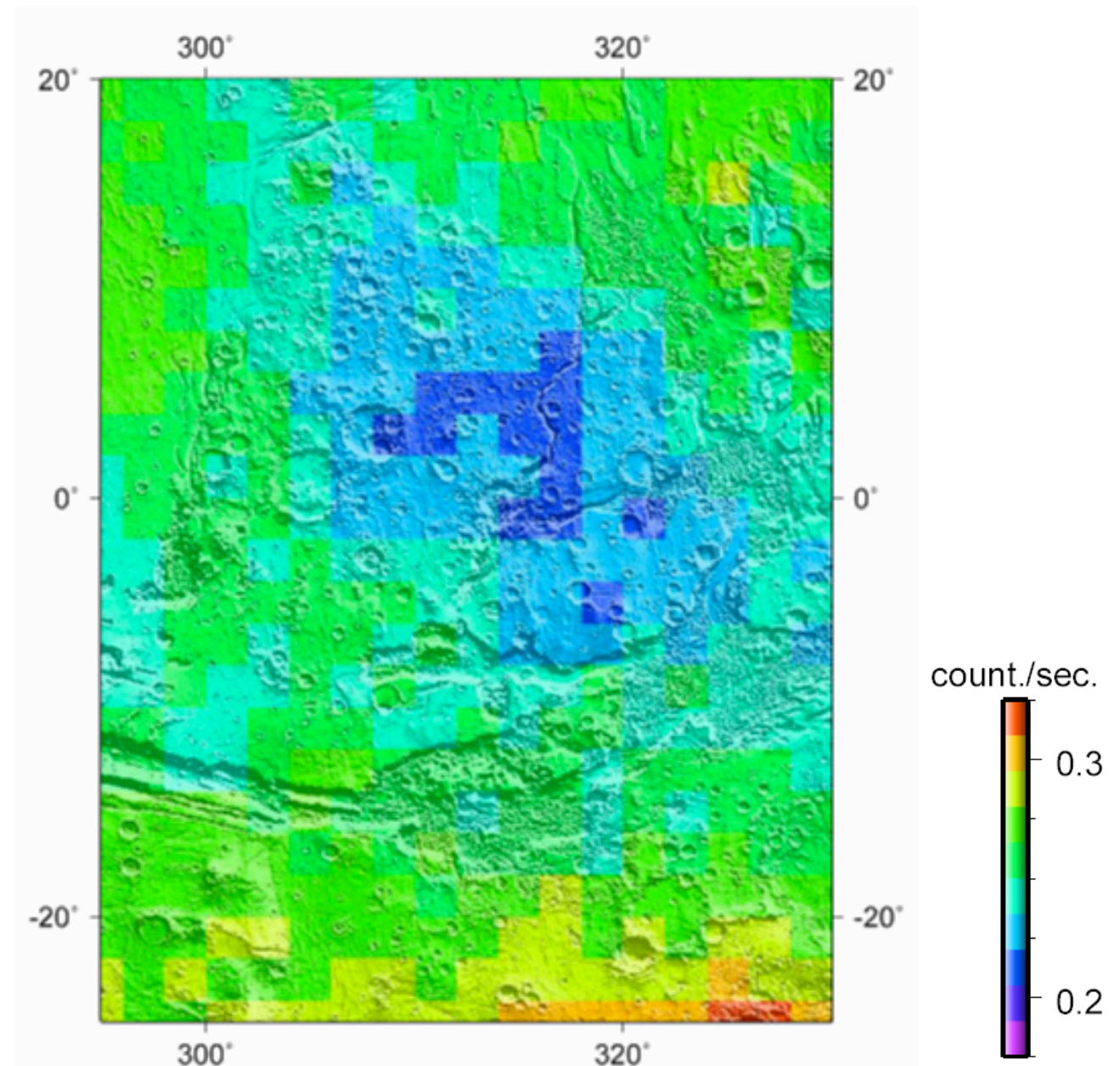


Marwth Vallis

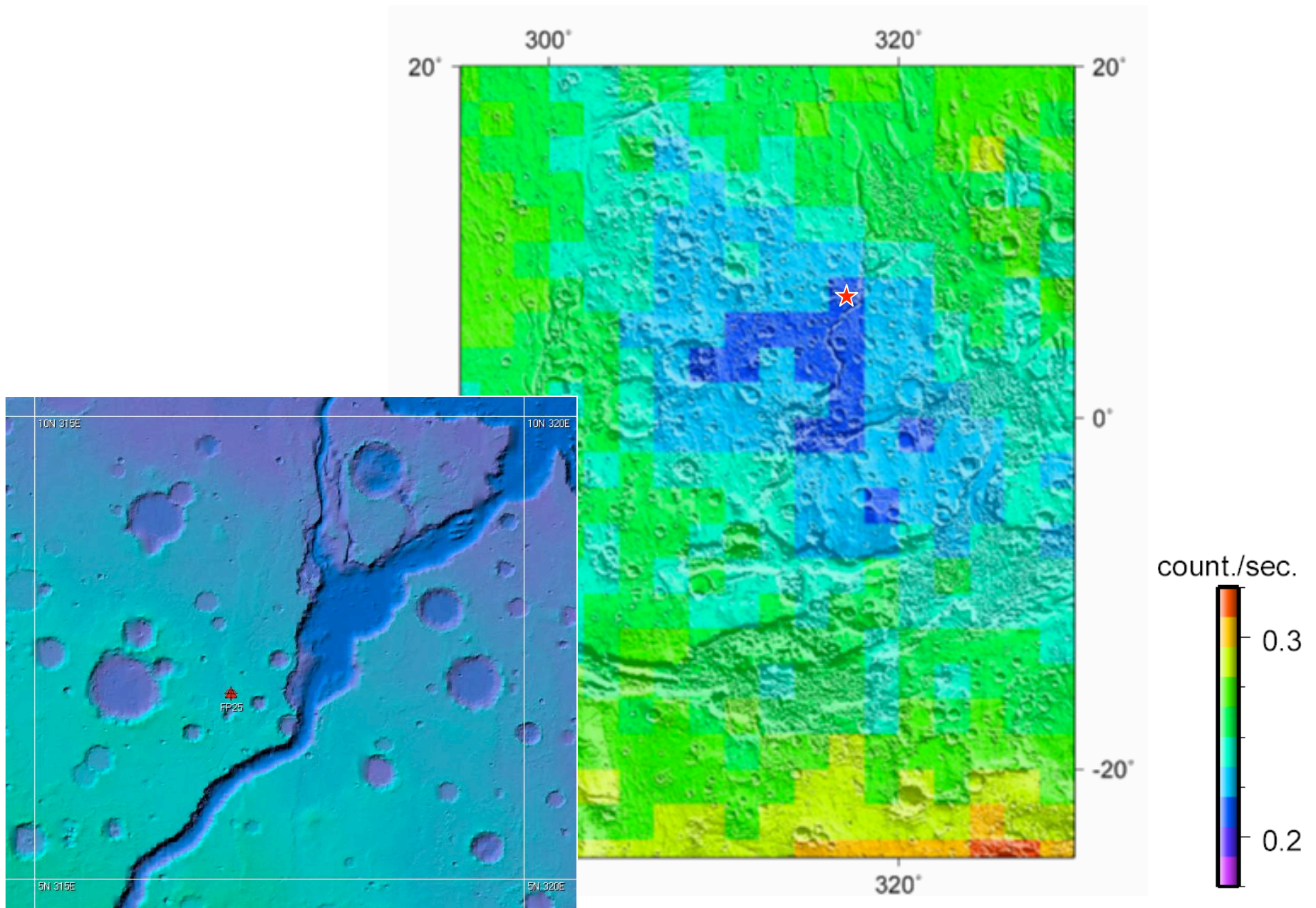


other proposed landing sites

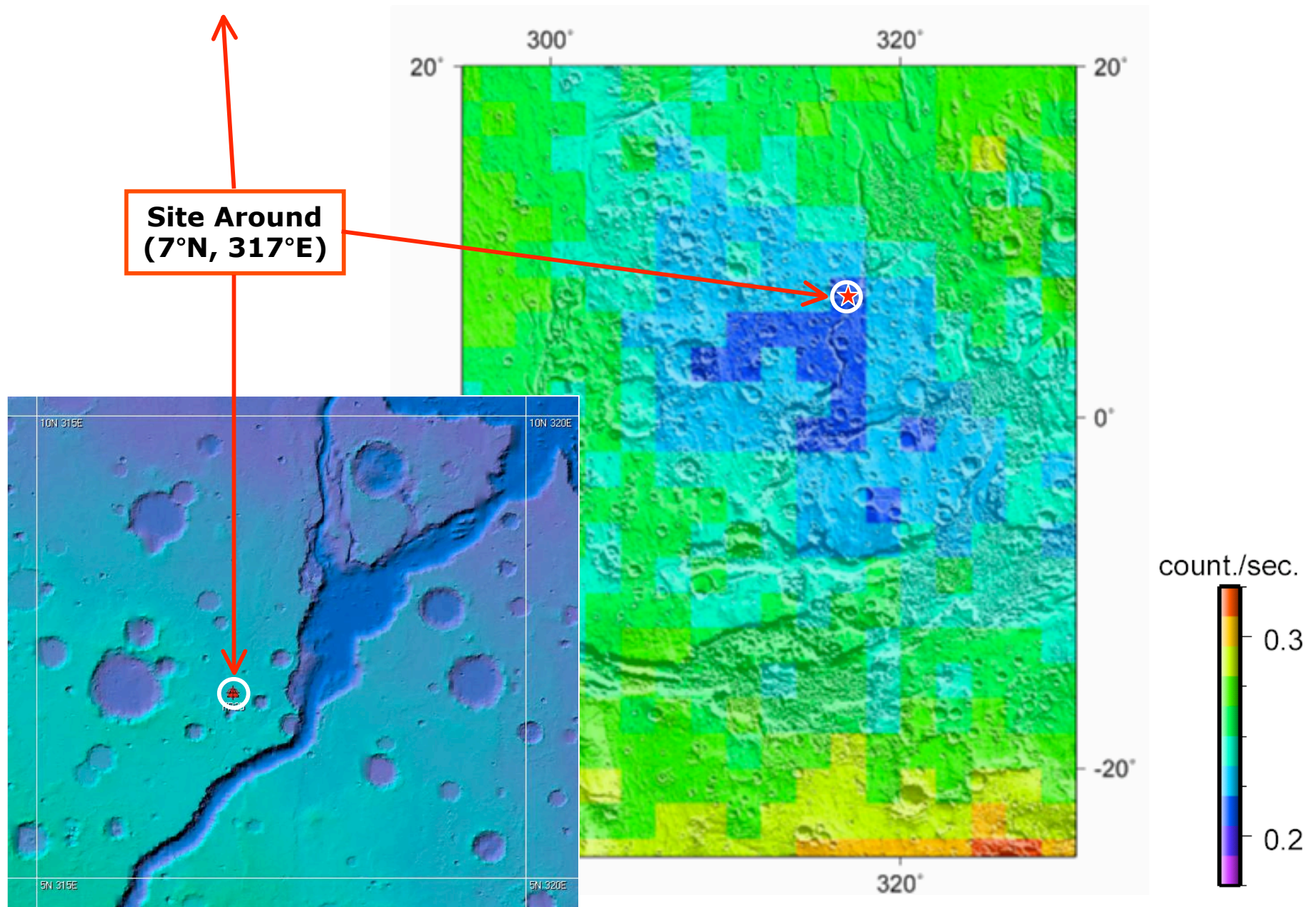
Shalbatana Vallis Regions (as current water site + Phillosilicate site)



Shalbatana Vallis Regions (as current water site + Phillosilicate site)



Shalbatana Vallis Regions (as current water site + *Phyllosilicate* site)



All four proposed landing sites satisfy the main engineering constraints,

... and more work is in progress

Region	Latitude	Elevation (km)	TES TI	Rocks (%)	Slope/2-5 km
Samara Vallis	- 24°	< -1	>200	5-10	<2°
Eos Chasma	- 13.4°	< -3	350 -420	10-18	1-2°
Aeolis fan-delta	- 5°	< -2.3	300 -350	7-12	<2°
Shalbatana Vallis	+ 7°				

Concluding Remarks:

(1) We believe that signatures of current water could be equally important for LS selection, as well as the signatures of past water activity

(2) We, as HEND/Odyssey + DAN/MSL team, want to be in the dialogue with the colleagues of Mars science community for optimal selection of MSL LS =

= that is why we present our candidates today

(3) Our candidates are not so much developed presently, as it would be necessary for successful participation in the current “five-the-best” selection process =

= but we will work much more within the next few months for addressing all known selection requirements, both scientific and engineering

= we need the mechanism to keep the opportunity for our candidates for consideration in future for potential selection