

NAI Mars Focus Group Videocon

Science and Landing Site Priorities for the Mars 2003 Mission

Presentations by:

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Discussion moderator: Jack Farmer (ASU)

Technical support: Maria Farmer, David Nelson & Leslee Unser (ASU) & Mike Fitzgerrall (NASA Ames)

Science Objectives for Mission

- Detailed description of the mission science objectives and requirements available at <http://athena.cornell.edu>
- Determine the aqueous, climatic, and geologic history of sites on Mars where conditions may have been favorable for the preservation of evidence of possible pre-biotic or biotic processes.
- Emphasis on locations possessing clear evidence for surface processes involving water.
- Wide range of possible settings, including former fluvial-lacustrine & hydrothermal environments.

**Recommendations of the Life
Subgroup of MEPAG**
(Mars Exploration Payload Assessment Group)

Life Subgroup GOAL:

DETERMINE IF LIFE EVER AROSE ON MARS

Objective 1: Determine if life exists *today*

Objective 2: Determine if life existed on Mars in the *past*

**Objective 3: Assess the extent of prebiotic organic
chemical evolution on Mars**

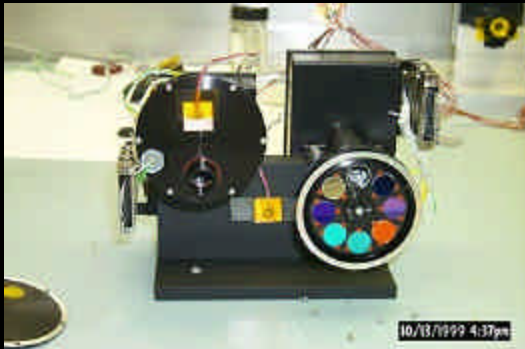
Prioritized MEPAG Investigations

- 1A. Map the present 3-dimensional distribution of water in all its forms.
- 1B. Determine the locations of sedimentary deposits formed by past surface and subsurface hydrological processes.
- 2. Carry out *in situ* exploration for possible liquid water in the subsurface.
- 3. Explore high priority candidate sites (i.e., those that provide access to near-surface liquid water) for evidence of extant (active or dormant) life.
- 4. Locate and sample aqueous rock samples for MSR to search for fossil biosignatures.
- 5A. Determine the array of potential energy sources to sustain biological processes (determine the distribution of potential energy sources for life (e.g. near-surface hydrothermal systems), redox state, distribution and abundance of biologically important elements (e.g. P and N).
- 5B. Determine the nature and inventory of organic carbon in representative soils and ices of the Martian crust.
- 5C. Search for complex organic molecules in rocks and soils.
- 6. Determine the distribution of oxidants and their correlation with organics.
- 7A. Determine the timing and duration of hydrologic activity during Martian history.
- 7B. Determine the changes in crustal and atmospheric inventories of organic carbon over time.

Mars Exploration Rover (MER)

- See <http://athena.cornell.edu>
- Twin rovers (MERs)
- Airbag landing system
- Surface operations
 - MERs are expected to operate on the surface of Mars for a minimum of 90 sols, with the second rover arriving 35 sols after the first
- Rover design and payload

'03 Rover & Payload



Pan Cam



Mini TES



APXS



Micro Imager



Mössbauer

Engineering Constraints

- For complete description of mission engineering constraints see:
 - <http://marsoweb.nas.nasa.gov/landingsites/mer2003>
 - <http://webgis.wr.usgs.gov/mer>
- Operation of rovers requires that MER-A be separated by $\sim 37^\circ$ from MER-B
- Both MERs must land below the -1.3 km MOLA defined elevation
- Power usage and thermal cycling restricts landing sites for MER-A to between 15°S and 5°N and MER-B to between 10°S and 10°N

Engineering Constraints (cont.)

- Landing error ellipses ~56 by 30 km for MER-A at 15°S and ~224 by 30 km for MER-B at 10°N
- Orientation of the ellipse rotates from 66° at 15°S to 87° at 5°N for MER-A and from 98° at 10°N to 81° at 10°S for MER-B
- Landing sites must appear hazard free at MOC scale and possess slopes <15°
- Total rock coverage should be less than 20% (based on thermal inertia) with <1% rocks being >0.5m high

Engineering Constraints (cont.)

- Radar reflectivity >0.05 ; extremely high albedo and low thermal inertia regions to be avoided
- Fine component thermal inertia values $>3-4 \times 10^{-3} \text{ cal cm}^{-2} \text{ s}^{-0.5} \text{ K}^{-1}$ or cgs units (equivalent to $125-165 \text{ J m}^{-2} \text{ s}^{-0.5} \text{ K}^{-1}$ or SI units)
- Low-altitude winds $<20 \text{ m/s}$

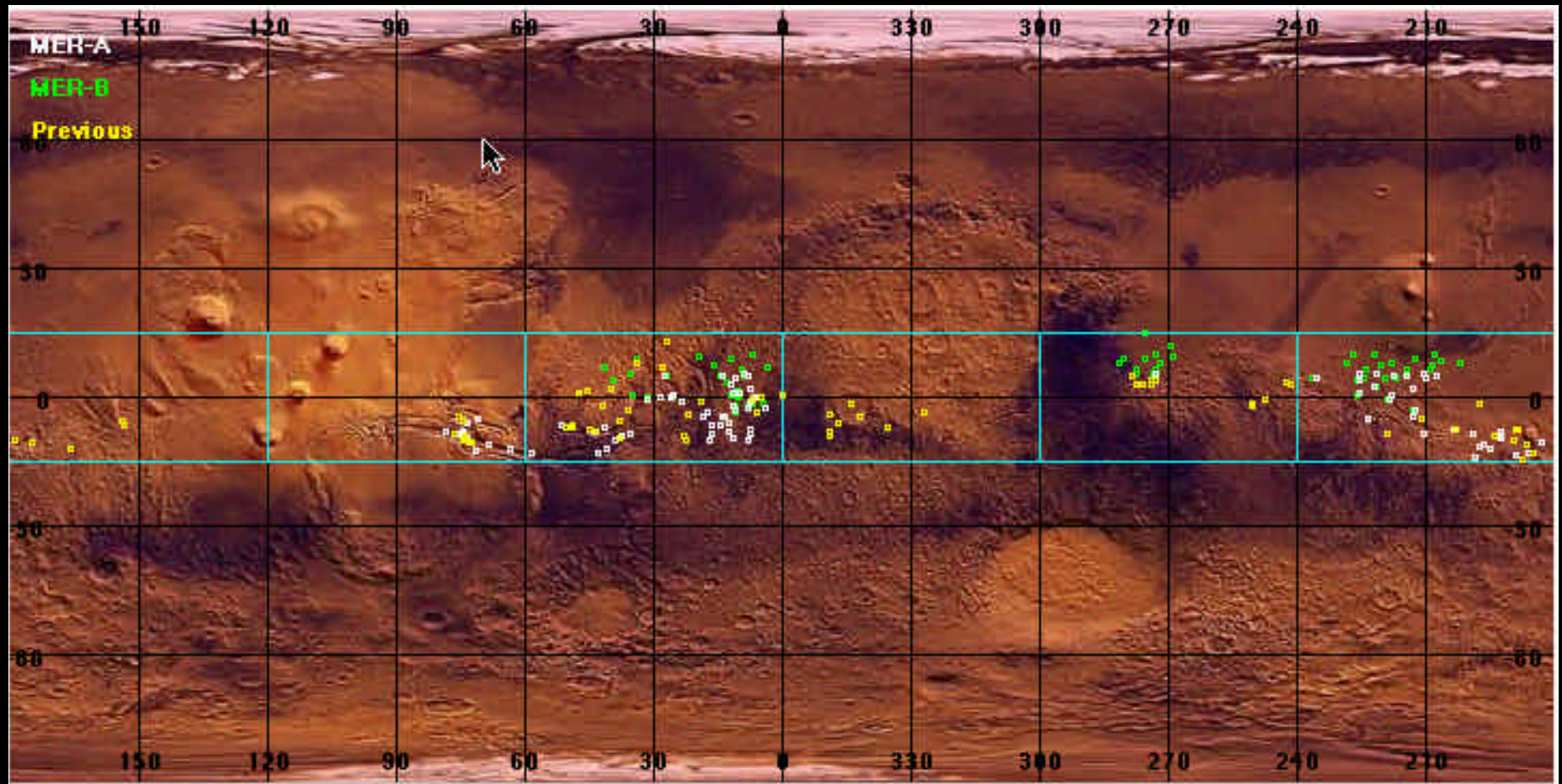
Potential Landing Sites

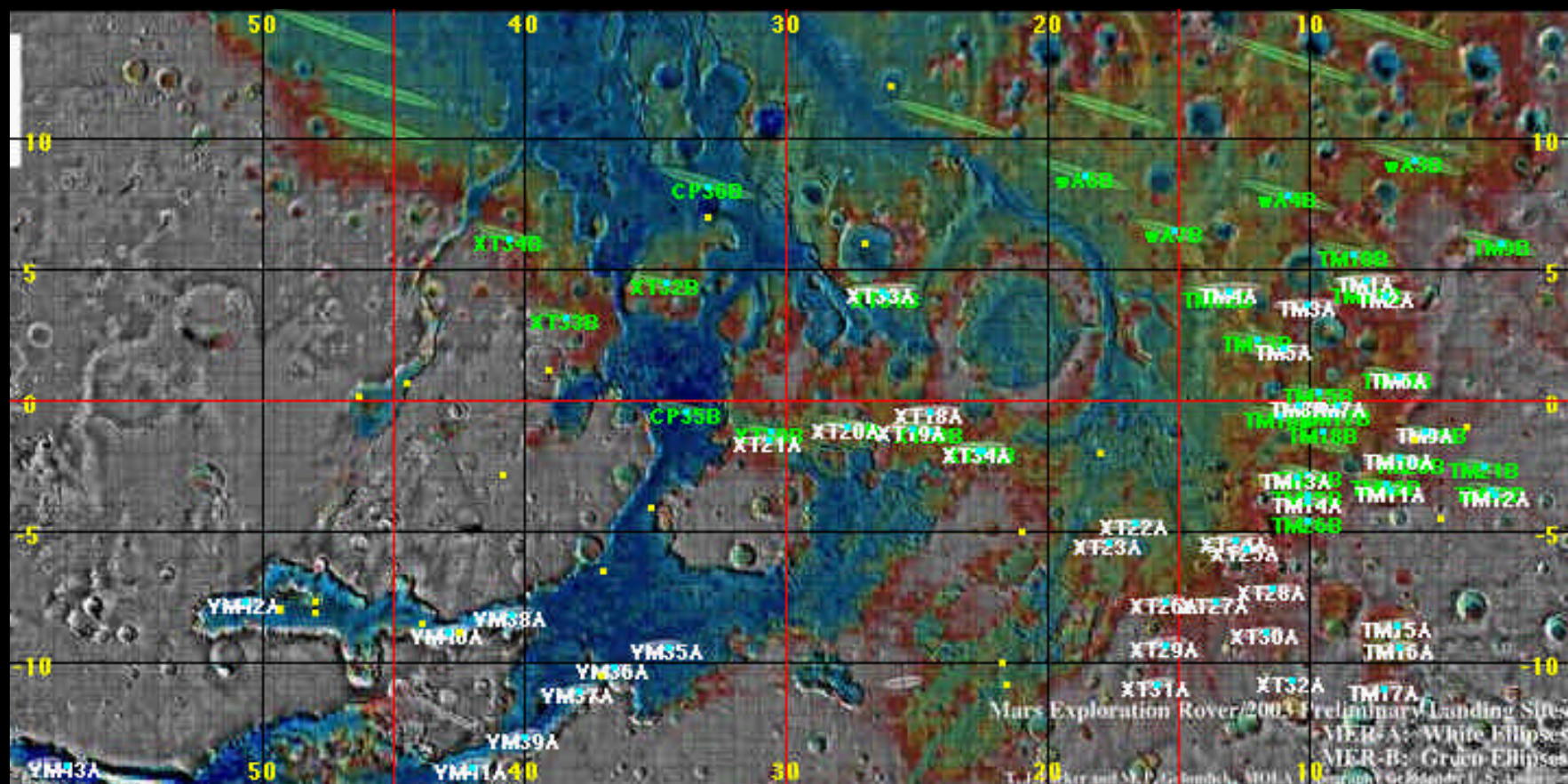
- Preliminary evaluation of potential landing sites was made for each 2.5° in latitude by placing ellipses of the proper size in all locations:
 - Below -1.3 km in elevation
 - With acceptable fine component thermal inertia values
 - Free of obvious hazards in the Mars Digital Image Mosaics (smooth and flat in the MDIM without scarps, large hills, depressions or large fresh craters)

Preliminary Evaluations

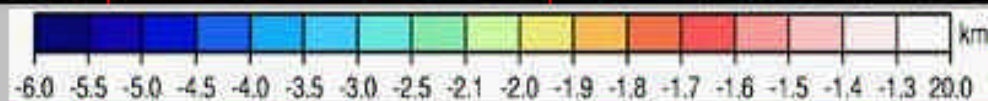
- ~185 potential landing sites shown to meet criteria (100 sites for MER-A & 85 for MER-B).
 - A complete listing of all of these sites can be viewed at:
 - <http://marsoweb.nas.nasa.gov/landingsites/mer2003>
 - <http://webgis.wr.usgs.gov/mer>

Potential Landing Sites





MOLA
(Golombek & Parker)



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Full Map

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Click in map to download:



Image



YRML

LANDING SITES:

☒ MER Sites

☒ MER Names

☒ Prior Proposals

Click map to download images & YRMLs

DISPLAY DATA:

☐ MDIMs

☐ Geology

☒ MOLA

or click-&-hold on landing sites for ID.

TES Thermal Inertia:

☐ interpolated

☐ uninterpolated

(Yellow dots are previously proposed sites from

TES Minerals:

☐ Andesite

☐ Basalt

☒ Hematite

the '98 and '99 Landing Site Workshops. Most

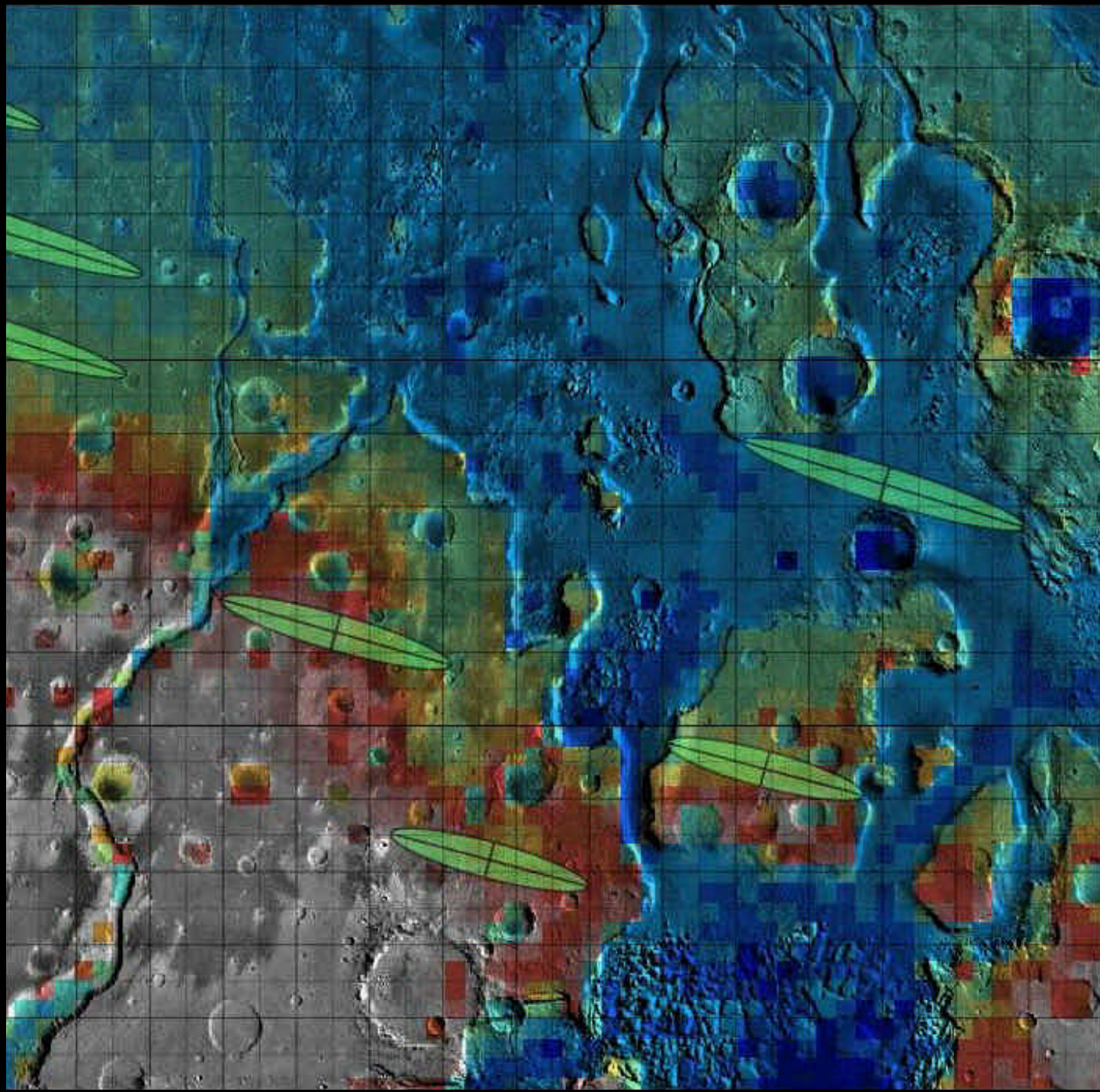
IRTM:

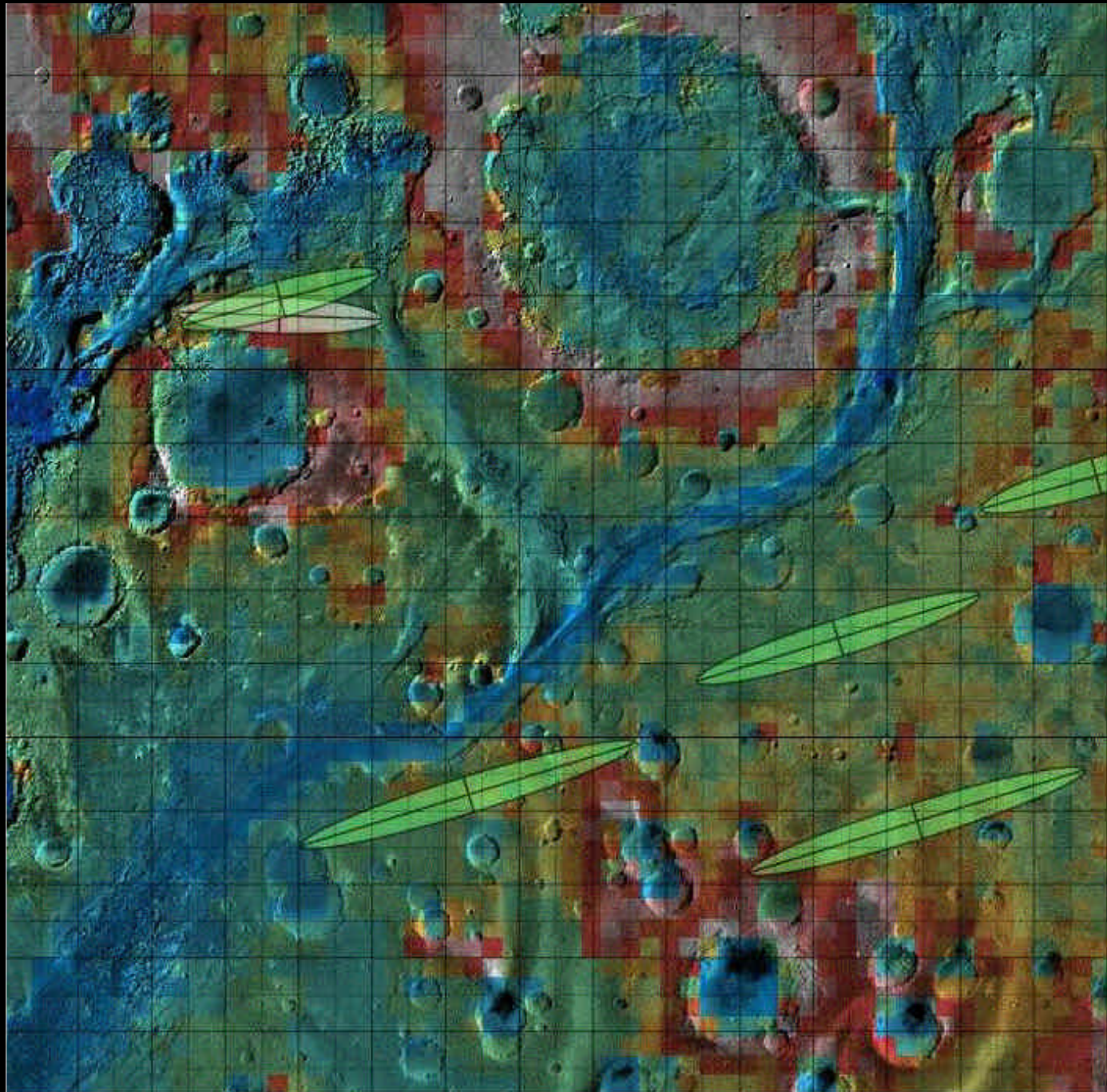
☐ Bulk Inertia

☐ FC Inertia

☐ Rock Abundance

can be clicked to access their abstracts.)







Geology

(Golombek & Parker)

Click geology legend below for full legend



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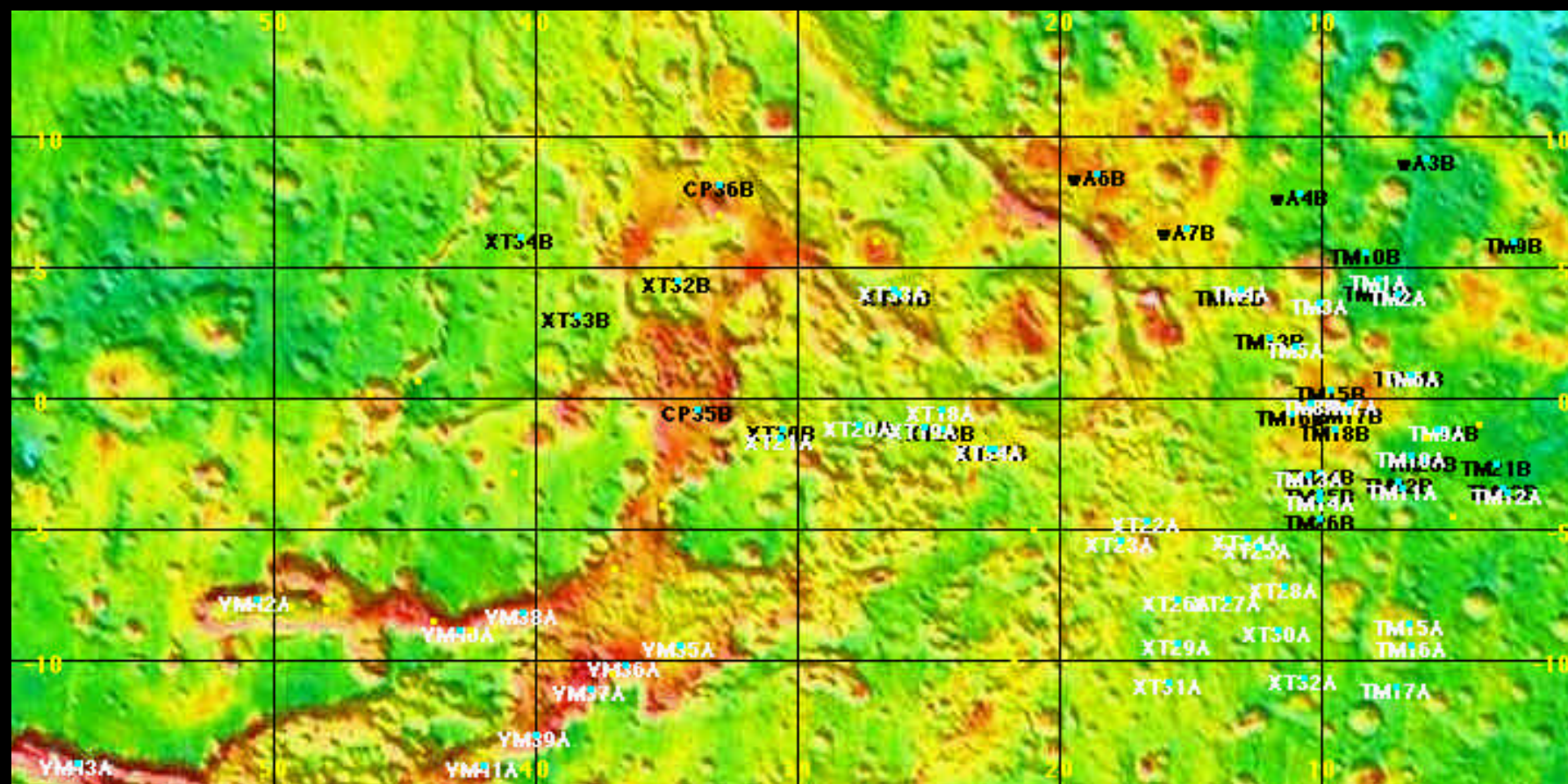
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TES Thermal Inertia

[interpolated]
(M. Mellon et al.)



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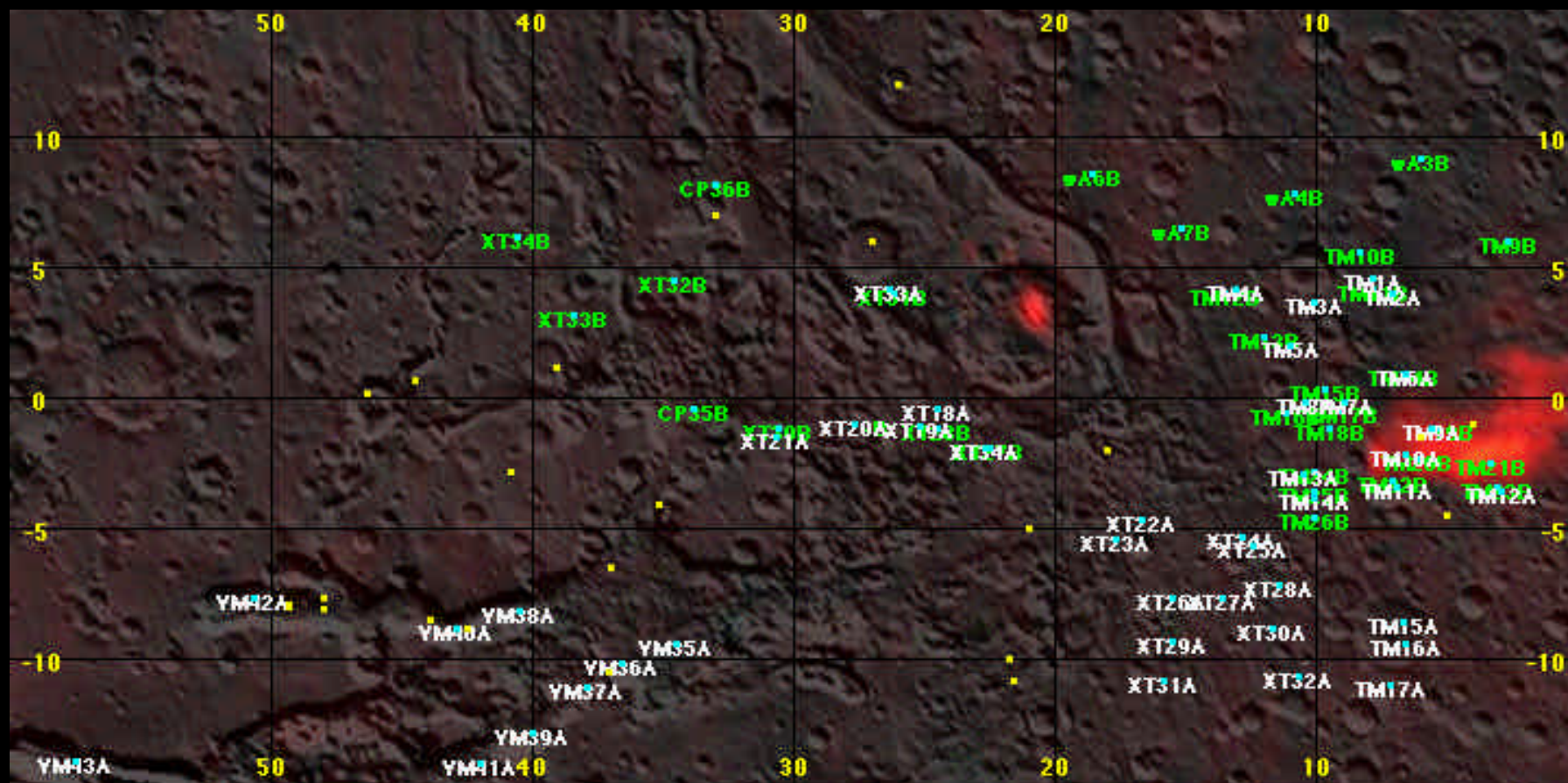
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TES Hematite
(J. Bandfield et al.)



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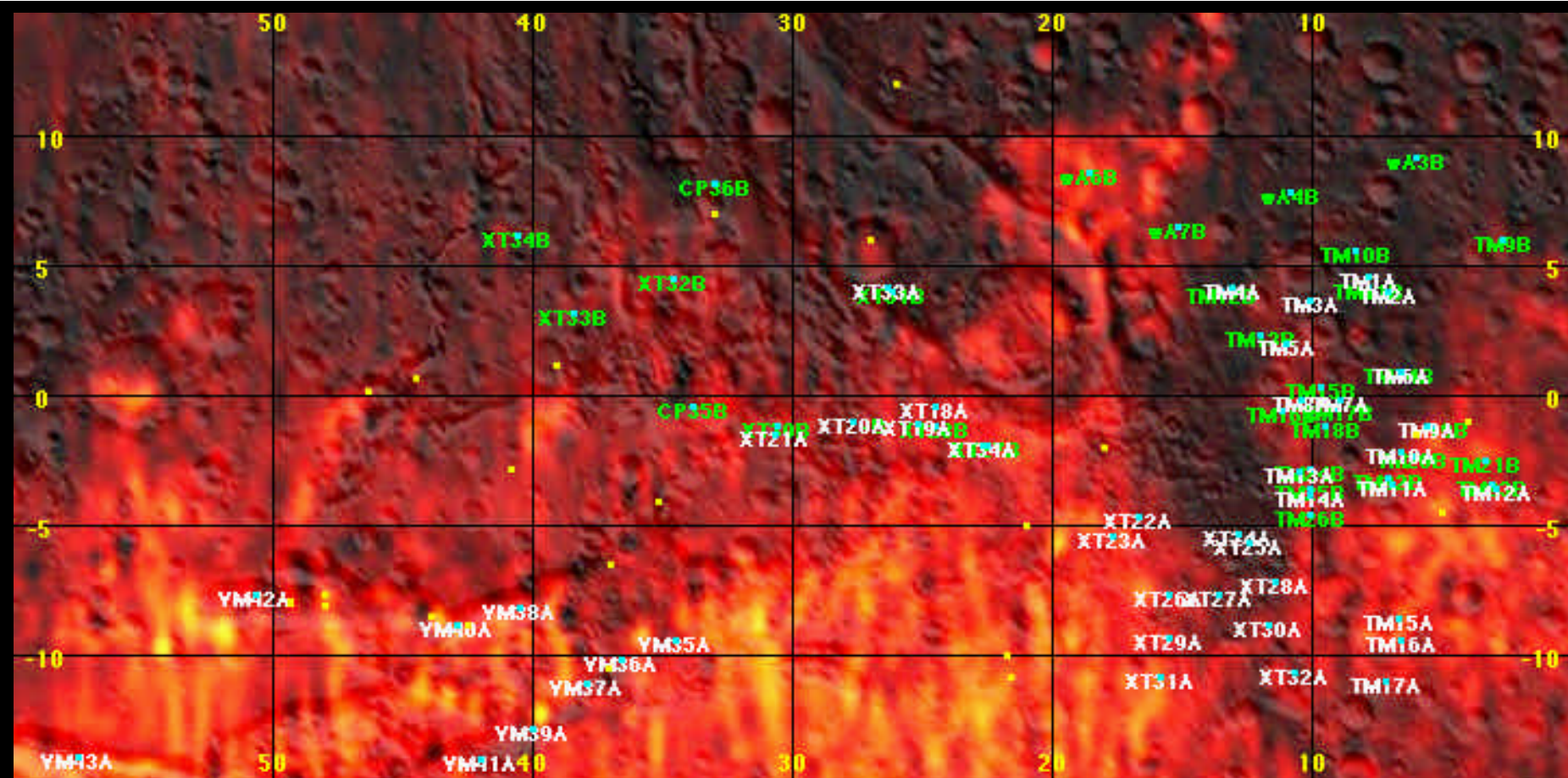
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TES Basalt
(J. Bandfield et al.)



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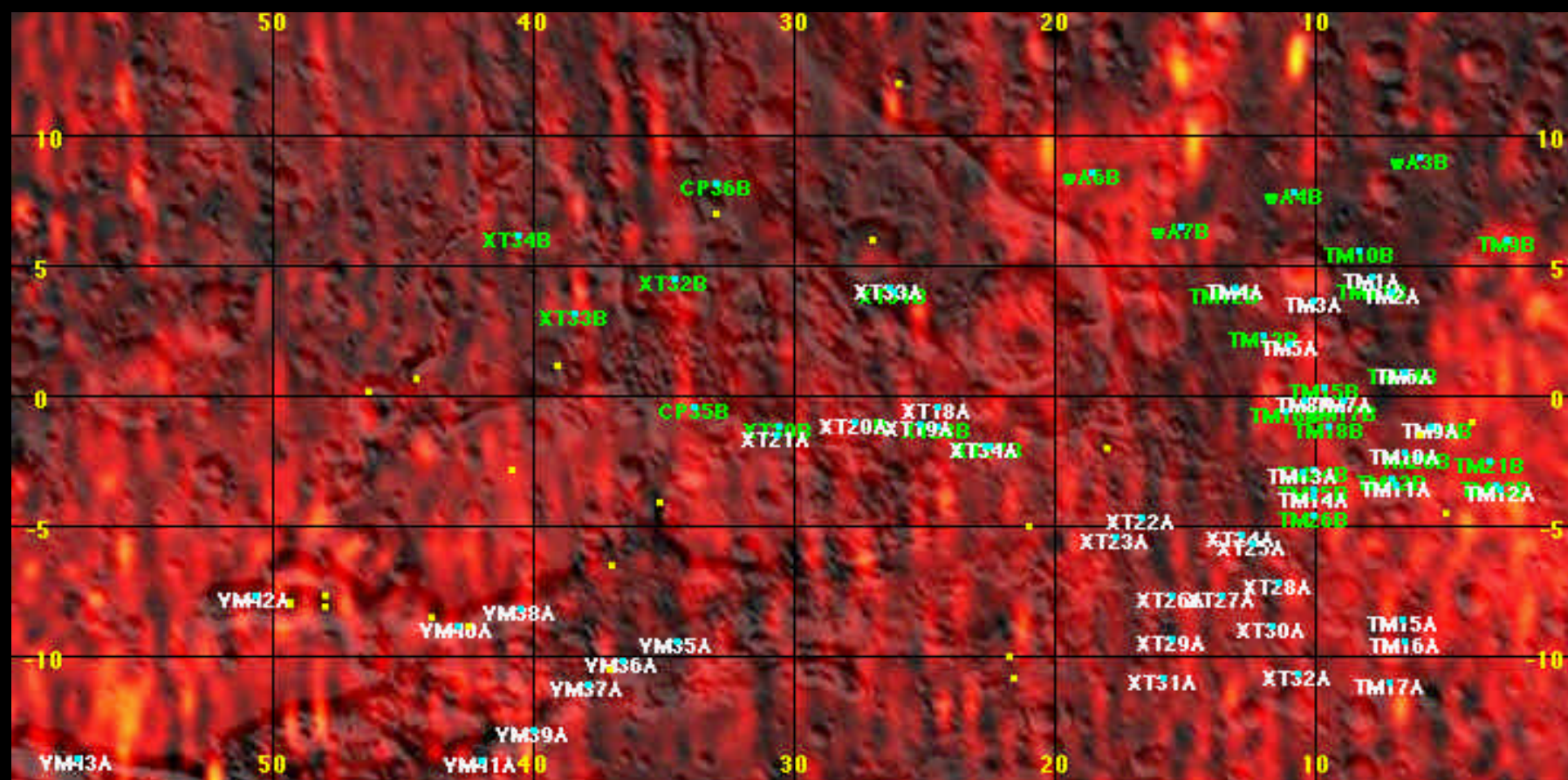
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Sites Reviewed In Today's Videocon:

- **Fluvial sites - Ron Greeley**

- Eos Chasma

- **Paleolake basins sites - Nathalie Cabrol**

- Gusev Crater

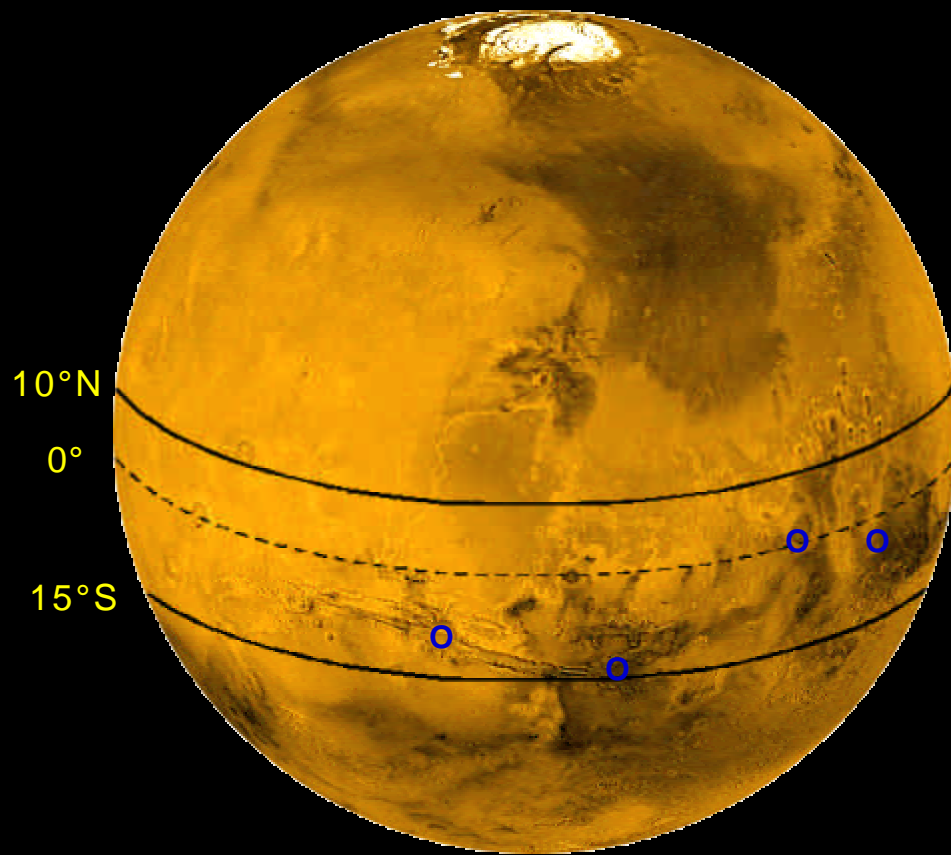
- Gale Crater

- **“Hematite” sites - Vicky Hamilton**

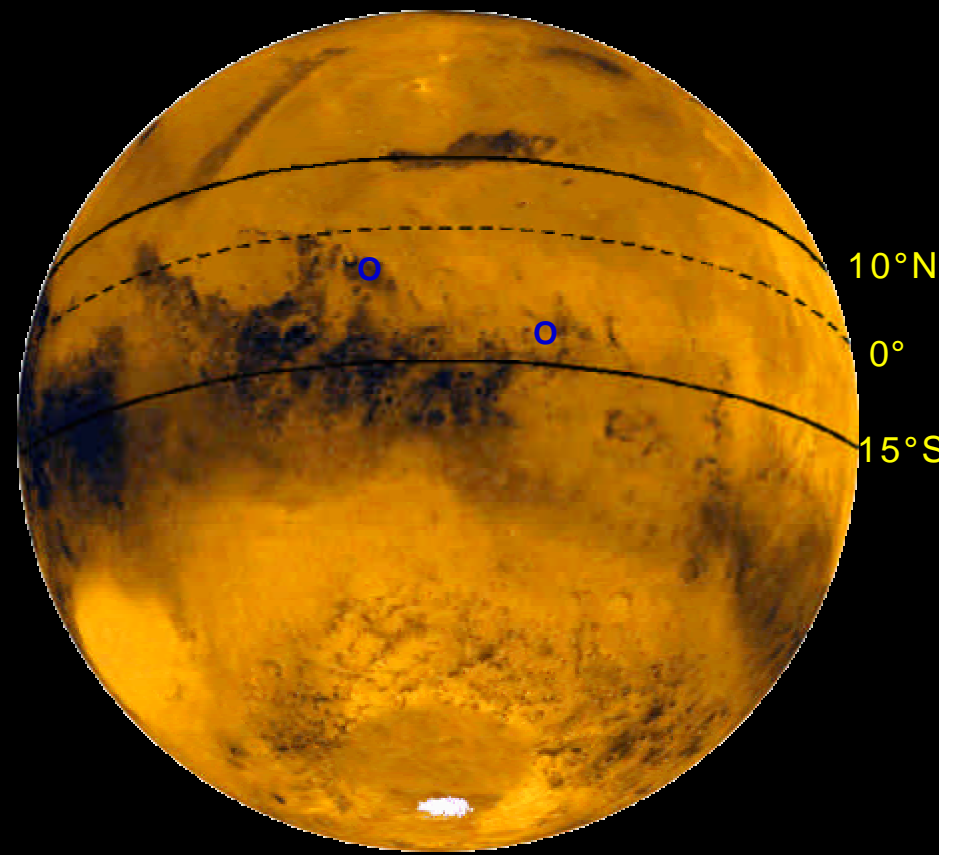
- Terra Meridiani

- Aram Chaos

Proposed Landing Sites for MER A & B



West Hemisphere
Centered at:
30°N, 30°W



East Hemisphere
Centered at:
30°N, 210°W

Where next?

