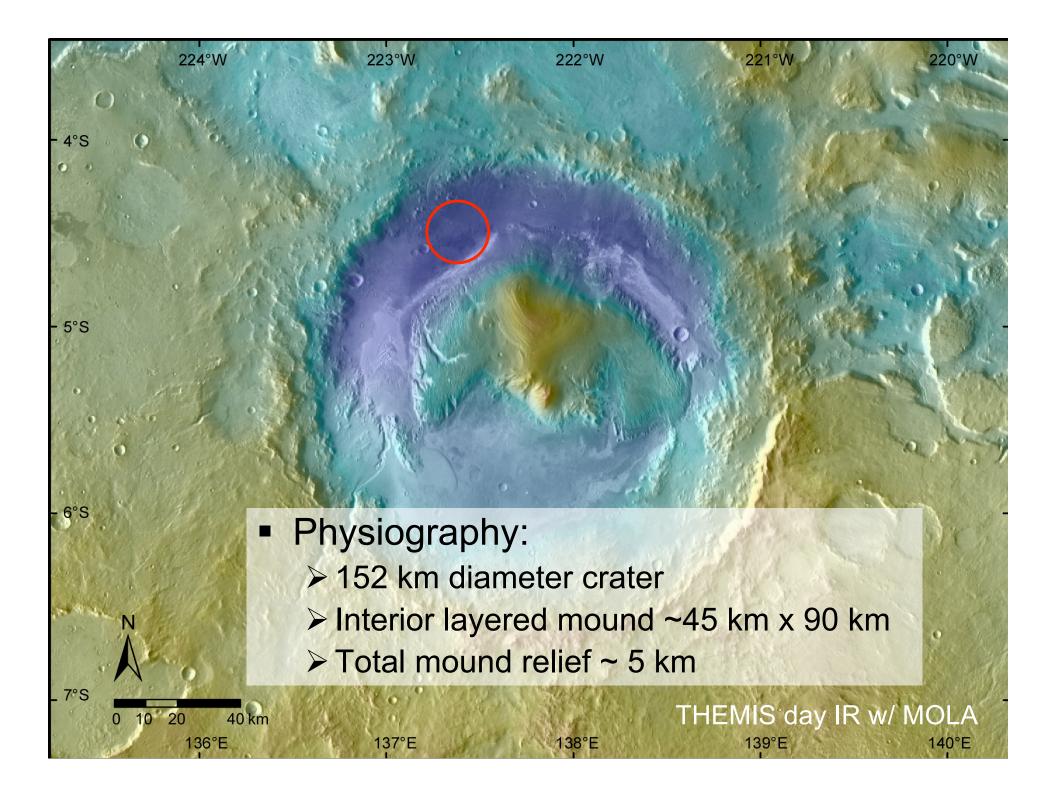
Gale Crater layered mound: A closed hydrologic system

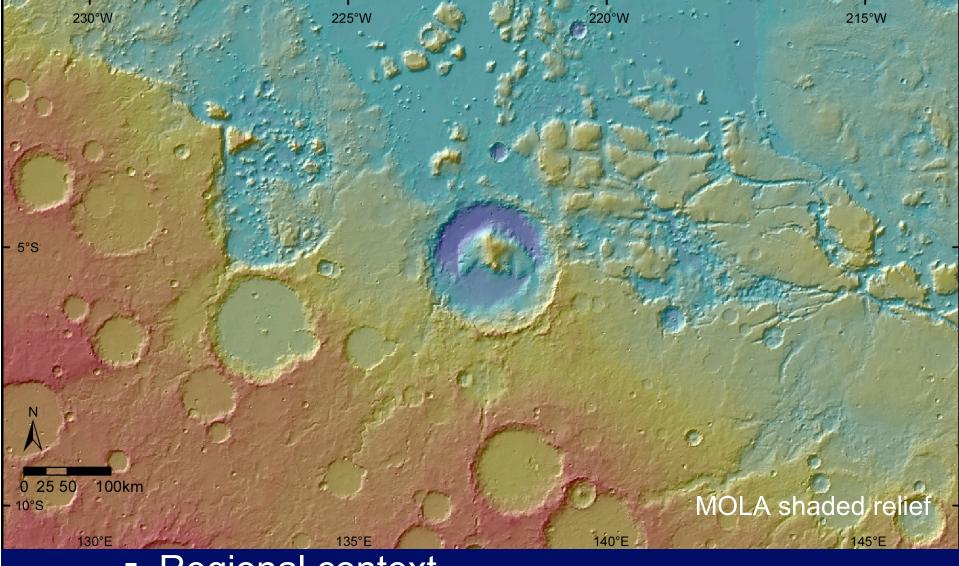
> Bradley Thomson<sup>1</sup> Nathan Bridges<sup>1</sup> James Bell<sup>2</sup> Ralph Milliken<sup>1</sup> Wendy Calvin<sup>3</sup>

<sup>1</sup>Jet Propulsion Lab, <sup>2</sup>Cornell University, <sup>3</sup>University of Nevada, Reno

# Outline

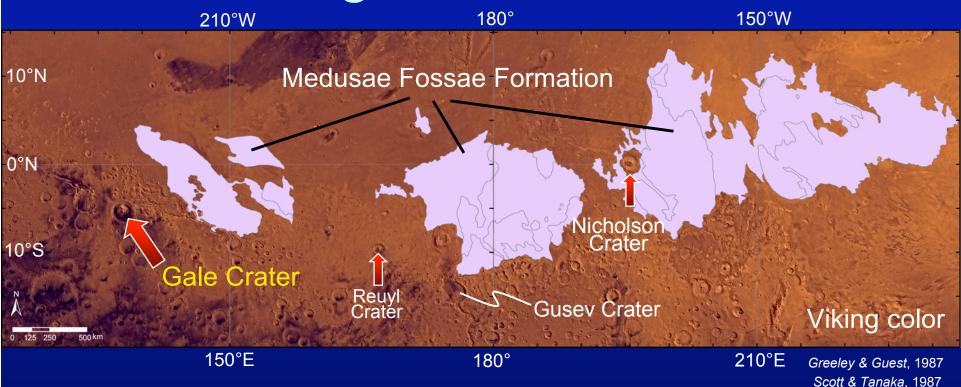
- 1. Regional context
- 2. Fluvial geomorphology
- 3. Nature of layered material
- 4. Layer compositional signatures
- 5. Inferred geologic history
- 6. Engineering constraints
- 7. Science objectives for MSL





- Regional context
  - Straddles hemispheric dichotomy boundary
  - Noachian plateau sequence
    - (1:15M scale geo units Npld, Npl2)

## **Regional context**



- Massive/layered sedimentary sequences not limited to Gale
- Medusae Fossae Formation (MFF) nearby
  Numerous additional outliers
- Also Gusev Crater, Apollinaris Patera in vicinity

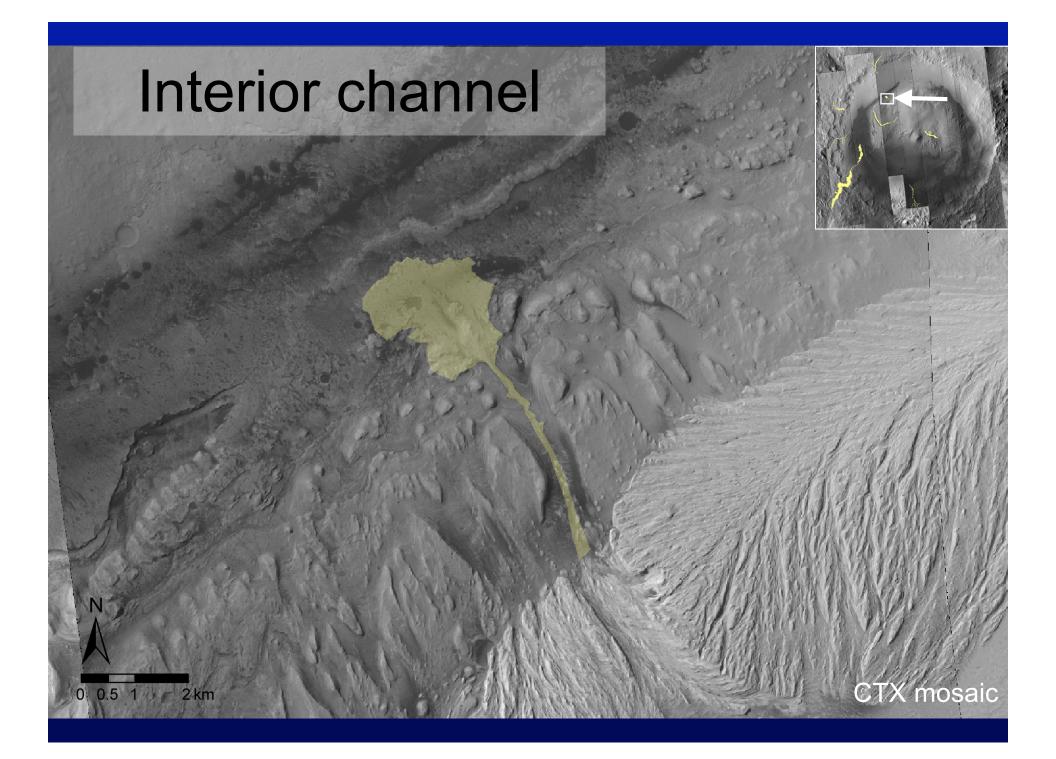
#### Interior channels on layered mound

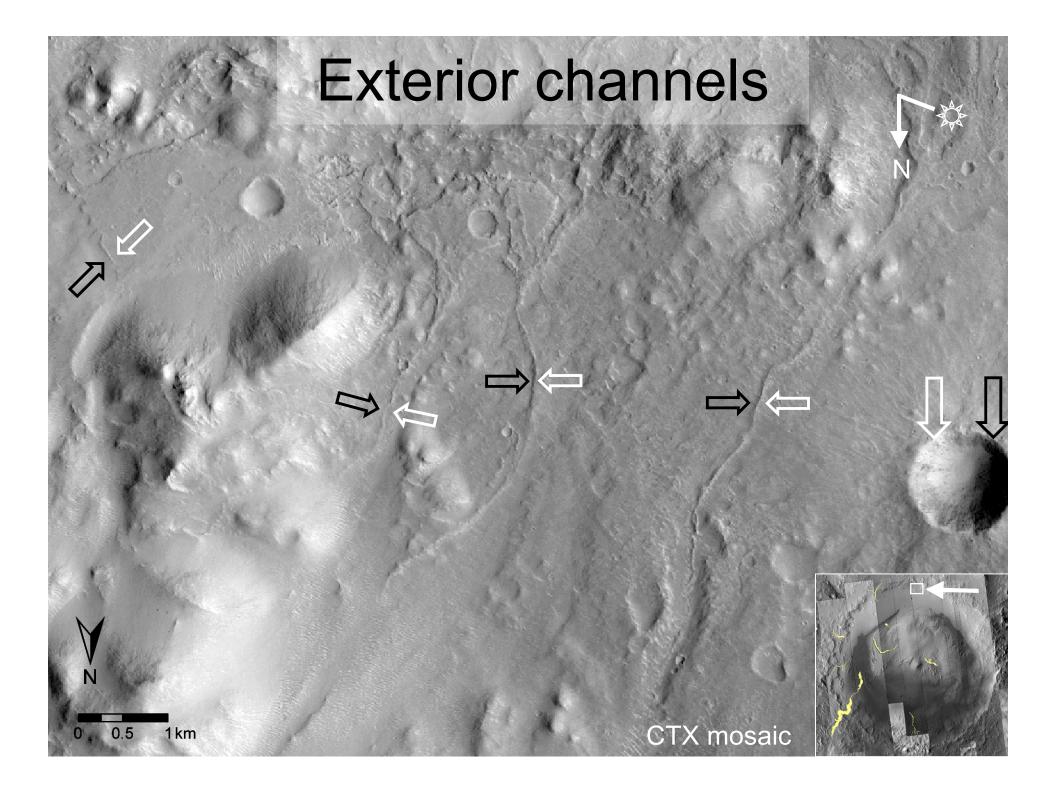
Fluvial activity

20km

10

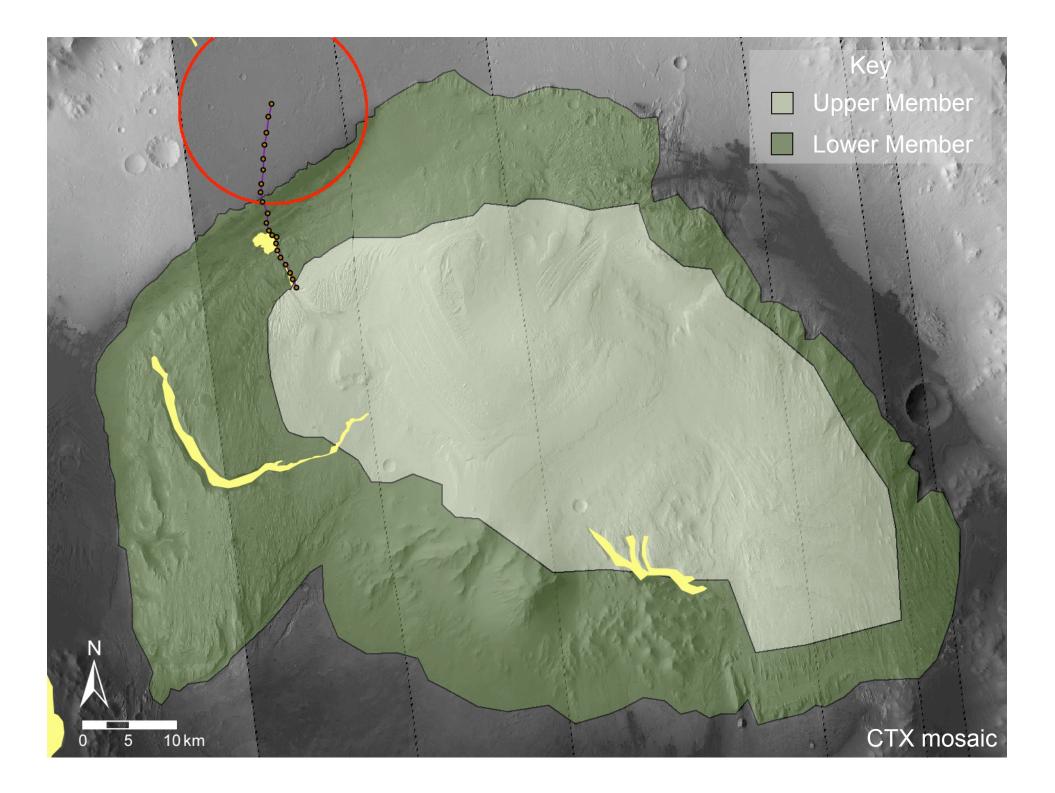
- Evidence for burial & exhumation of fluvial channels
- Source, sink, and transportation pathway preserved
- Exterior channels draining inward dissecting crater rim

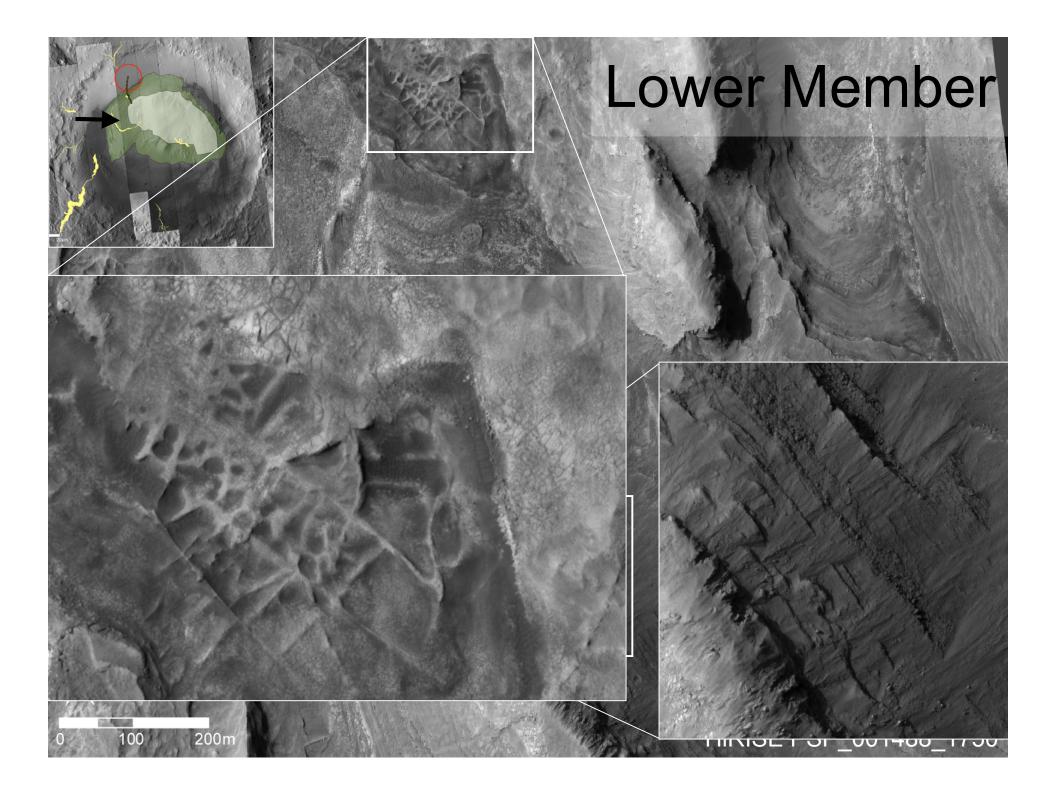


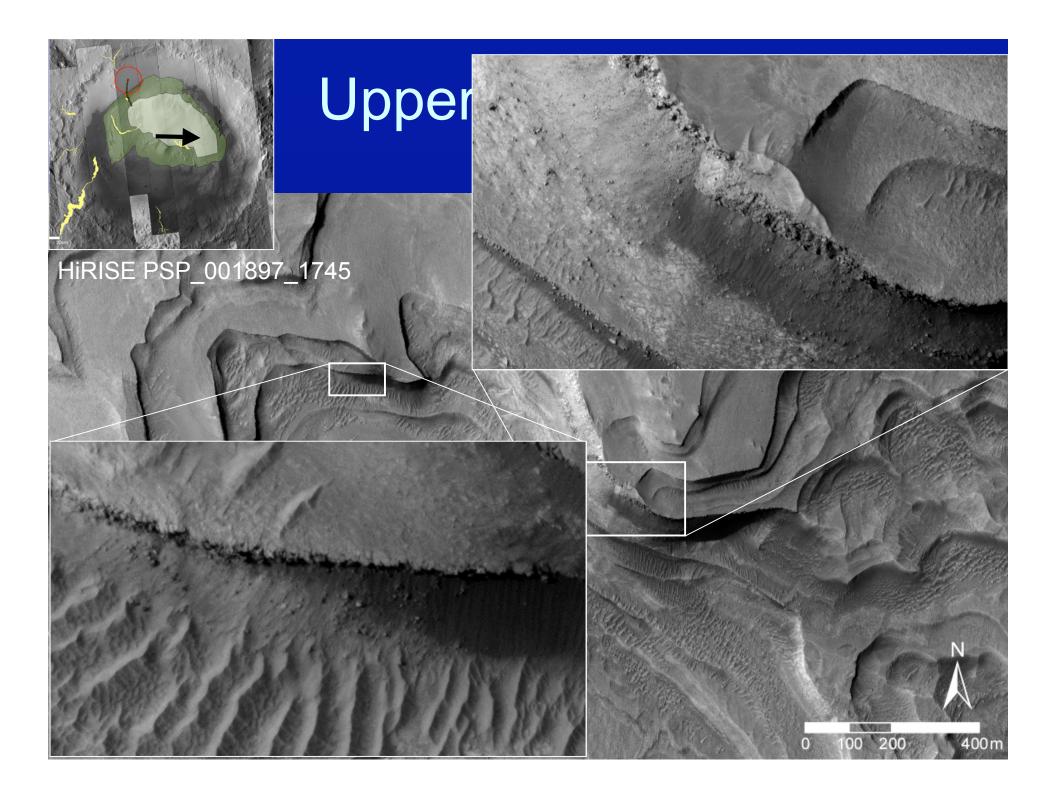


## Nature of layered material

- Mound consists of two distinct members
- Lower member:
  - Finely layered
  - Conformal contact relationships
  - Total thickness ~1.5 km
- Upper member:
  - Erosional contact with lower member [Edgett & Malin, 2000]
  - Both massive and layered units
  - Max thickness ~3.5 km (mean 2.5 km)
- Stratigraphy records significant changes in depositional and erosional regimes

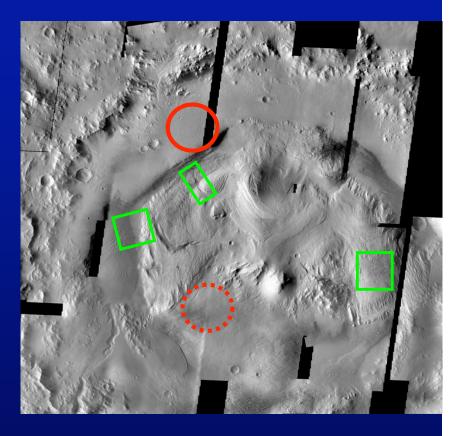






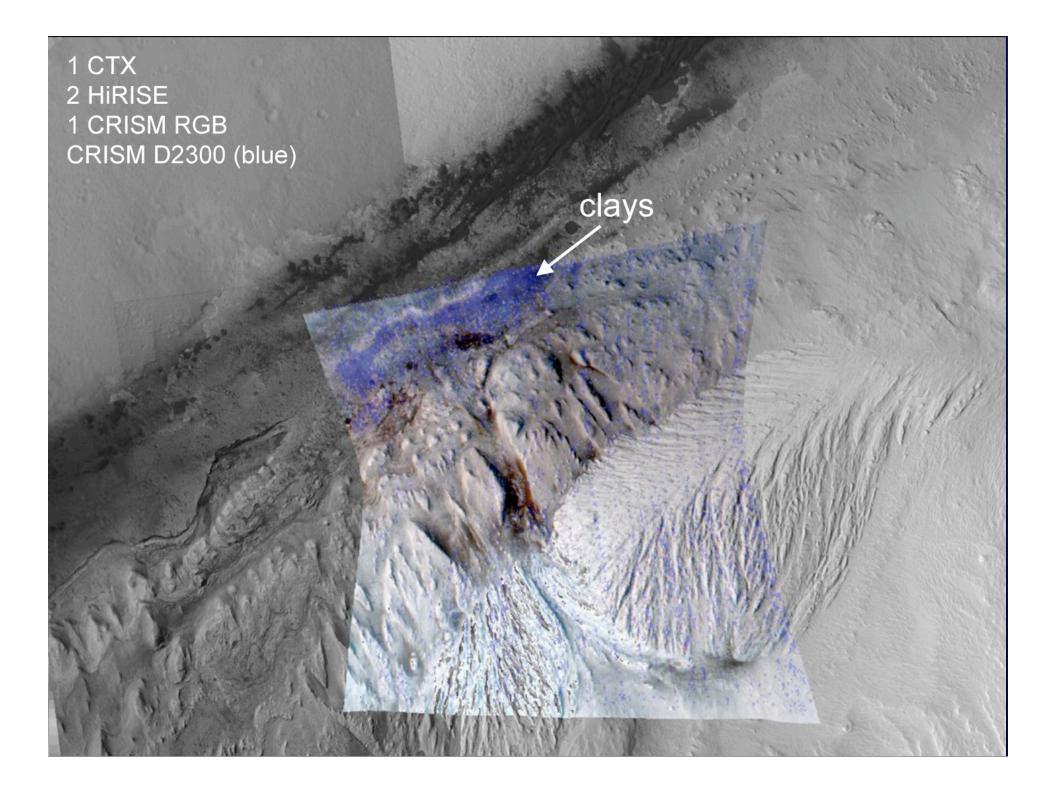
## Layer compositional signatures

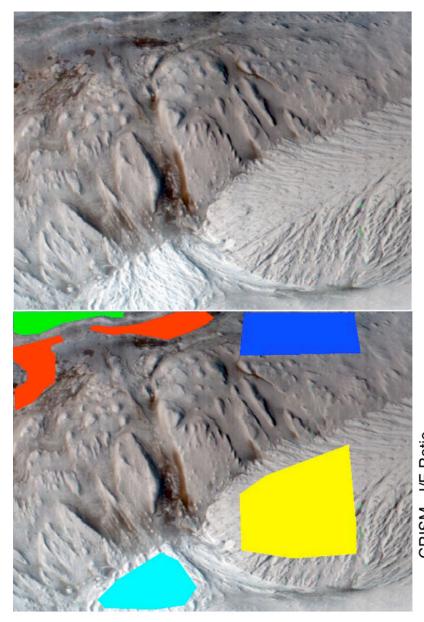
- Evidence for phyllosilicates
- Alteration mineralogy
- Implies large water/rock ratios, moderate to high degrees of alteration



Red circle: potential landing ellipse, dashed = backup

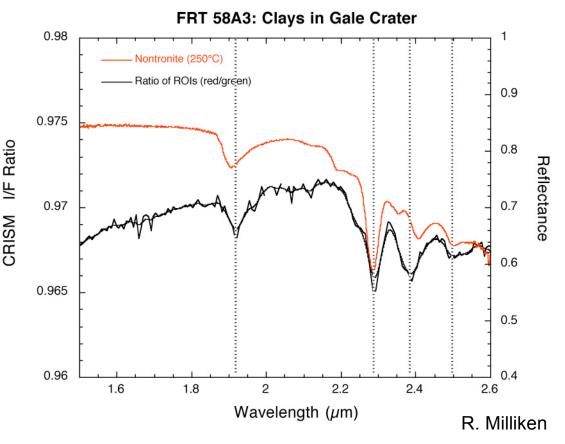
Green boxes: CRISM FRT locations (approx)





Used Region of Interests (ROIs) to get higher S/N and improved spectral ratios; best results are achieved when using ROI's at similar elevations. Absorptions are most consistent with an Fe/Mg-bearing smectite. There are clear spectral features at 1.91, 2.28, 2.39, and 2.5  $\mu$ m.

Similar to partially dehydrated nontronite, although the 1.9 and 2.4 µm bands are shifted slightly.



# Inferred geologic history

- Superposition indicates Gale Crater predates fretted terrain formation (dichotomy boundary)
- Lower Member of layered mound formation
  - Possible lacustrine or distal fluvio-deltaic contributions
  - Mineralogic evidence for aqueous alteration
  - Mass wasting contribution from degraded northern rim
- Upper Member of layered mound formation
  - Erosional contact with Lower Member indicates depositional hiatus
  - Massive to finely layered units, low thermal inertia => suggests airfall deposit possibly linked to Medusae Fossae Formation
  - Intermittent fluvial activity possible volatile component mixed with dust/ash component
- Continued eolian erosion, some potential late-stage fluvial activity

# Engineering constraints

Engineering Parameter	Requirement	Gale site value
Latitude	45°N to 45°S	4.5°S
Elevation	≤ +1 km	-4.4 ± 0.1 km
Slopes (2-10 km length scale)	≤ 20°	2.78 km baseline: 6.7° max
Terrain Relief (1-2 km length scale)	≤ 43 m relief at 1 km; ≤ 720 m at 2 km	1.39 km: 95 m relief 2.78 km: 328 m relief
Terrain Relief (0.2-1 km length scale)	≤ 43 m relief	T.B.D.
Slopes (2-5 m length scale)	≤ 15°	T.B.D. by photoclinometry

# Engineering constraints

Engineering Parameter	Requirement	Gale site value
Rock height	$\leq$ 0.55 m (< 0.50% chance of 0.55 m rock in 4 m <sup>2</sup> ). Suggests low to moderate rock abundance.	IRTM rock abundance: 10%
Radar reflectivity	Ka band radar backscatter cross-section (> -20 dB and < 15 dB)	T.B.D.
Load bearing surface	Thermal inertia >100 J/m <sup>2</sup> s <sup>0.5</sup> K albedo <0.25 radar reflectivity >0.01	TES TI: 483 TES albedo: 0.238 TES DCI: 0.96

## Science objectives

- Confirm nature of alteration assemblage
  - What minerals are present?
    - What are grain sizes, shapes, textures, relation to other constituents?
  - Assess degree of alteration, biologic habitability potential
- Determine origin of lower and upper sedimentary member formation:
  - Subaerial (airfall, impact, mass wasting) and/or subaqueous (lacustrine, fluvial)?
- Explore closed hydrologic system
  - Sources, sinks, and transportation pathways all accessible in single locality

# MSL instrument suite

Acronym	Full Name	Description	
MastCam	Mast Camera	Mast-mounted stereo camera	
MAHLI	Mars Hand Lens Imager	Arm-mounted surface imager	
APXS	Alpha Particle X-Ray Spectrometer	Arm-mounted chemistry probe	
ChemCam	Chemistry & Camera	Laser induced breakdown spectroscopy with remote micro-imager	
CheMin	Chemistry & Mineralogy	XRD / XRF delivered sample analysis	
SAM	Sample Analysis at Mars	GCMS & Tunable Laser Spectrometer for delivered sample analysis	

# Science Traceability Matrix

Goal	Objective	Instruments
Confirm nature of alteration assemblage	Identify alteration minerals	APXS, CheMin, SAM, ChemCam
	Characterize grain sizes, shapes, textures, relation to other constituents	MAHLI, ChemCam
	Assess degree of alteration, biologic habitability potential	APXS, CheMin, SAM, ChemCam
Determine layer origin(s)	Examine constituent particles: grain size, sorting, shapes, textures	MastCam, MAHLI, ChemCam
	Look for bedding relationships, sedimentary structures	MastCam, ChemCam
	Identify primary and secondary mineralogy	APXS, CheMin, SAM, ChemCam
Explore closed hydrologic system	Determine cementation agent, evidence of alteration	all



### Additional considerations

- Unlike Meridiani, here the entire stratigraphic column is exposed
- Any rover traverse path will start with the oldest units
  - Lower strata appear to be clay-bearing
  - Can drive up section to access younger strata
- Key alteration mineralogy can be accessed very early in mission
- Low energy depositional environment has high biomarker preservation potential

