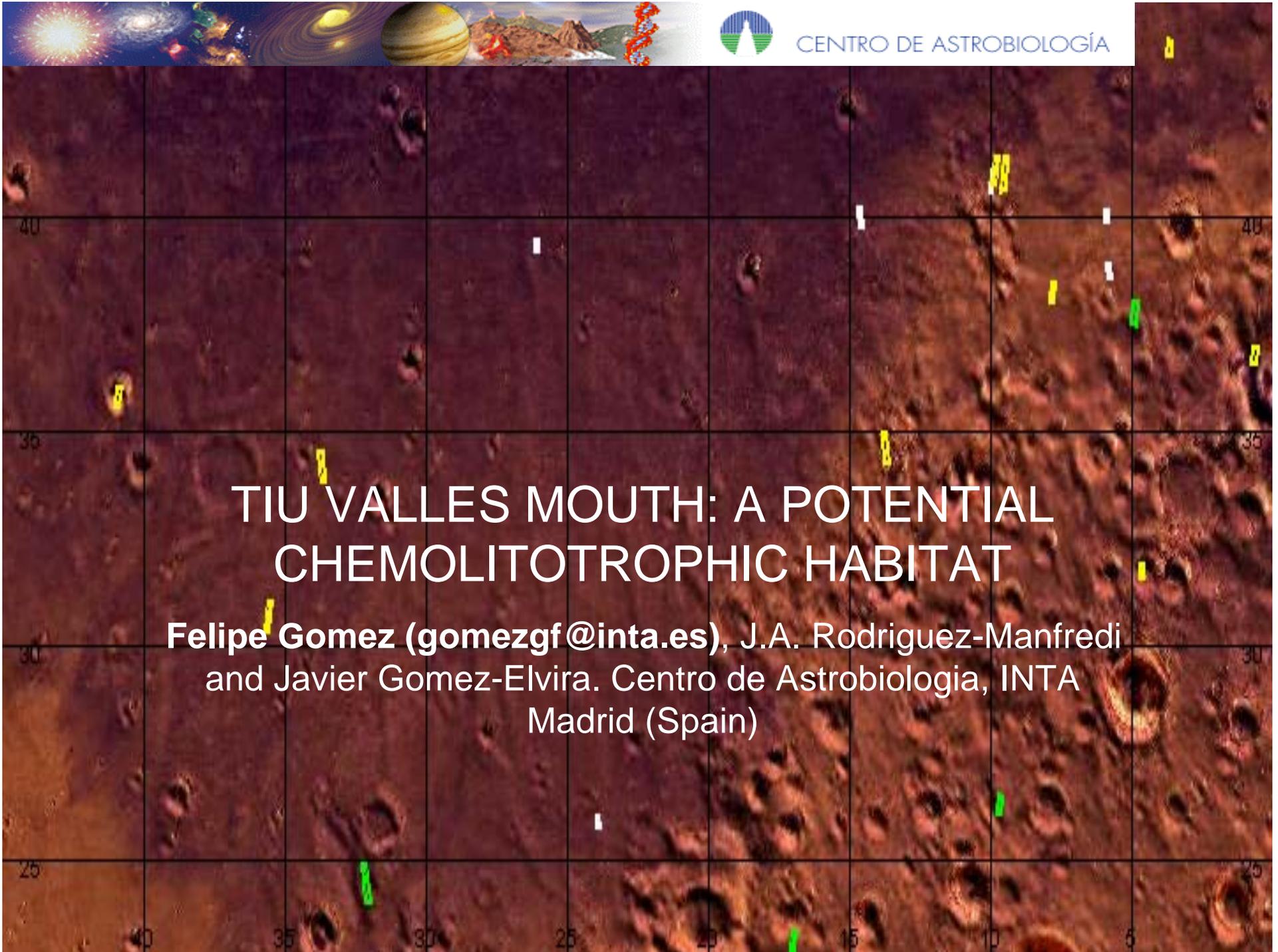
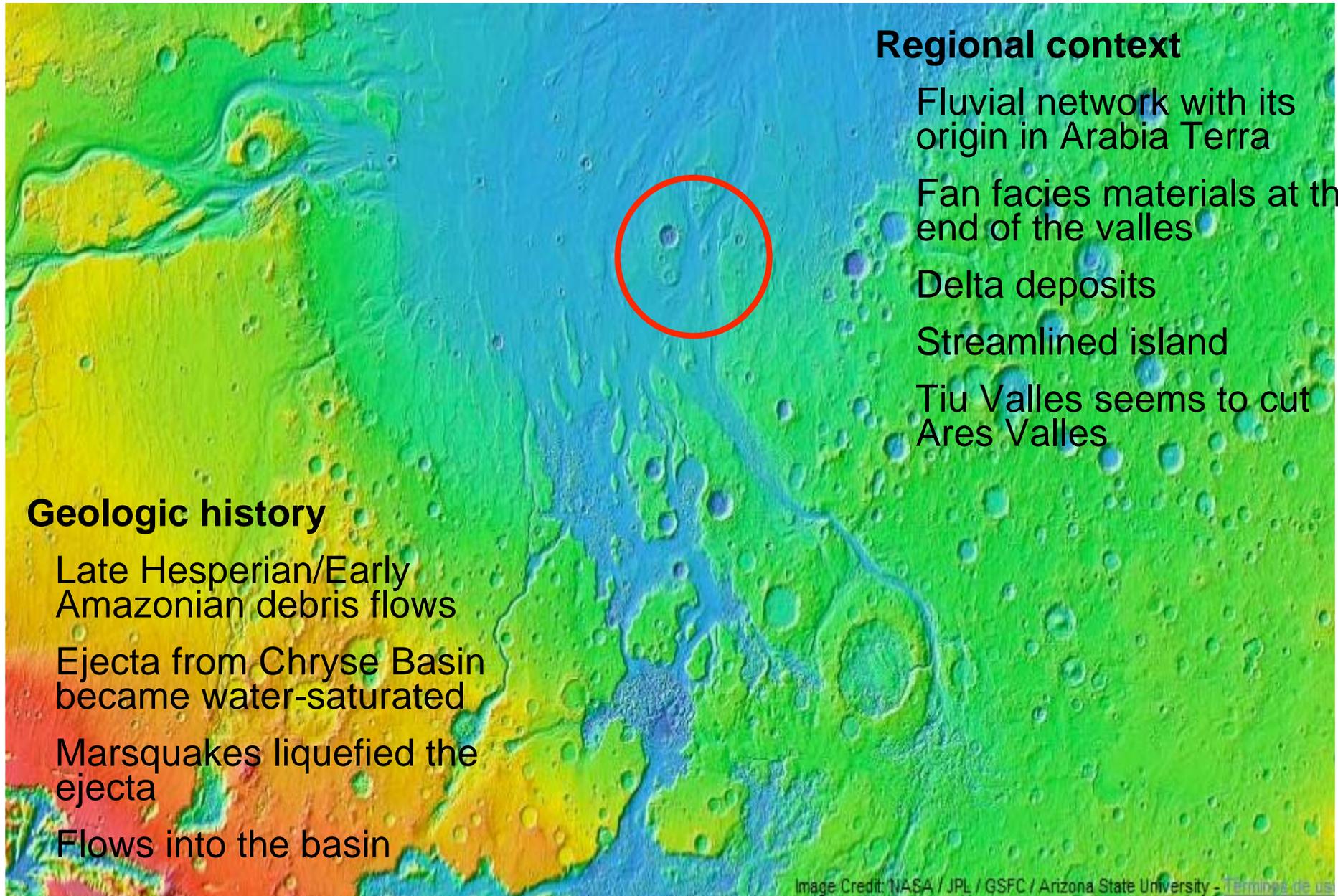
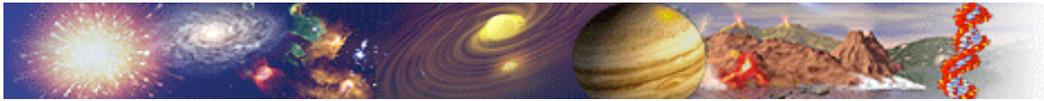




TIU VALLES MOUTH: A POTENTIAL CHEMOLITOTROPHIC HABITAT

Felipe Gomez (gomezgf@inta.es), J.A. Rodriguez-Manfredi
and Javier Gomez-Elvira. Centro de Astrobiologia, INTA
Madrid (Spain)



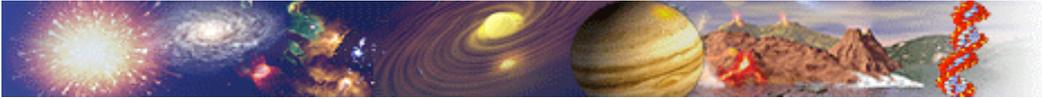


Regional context

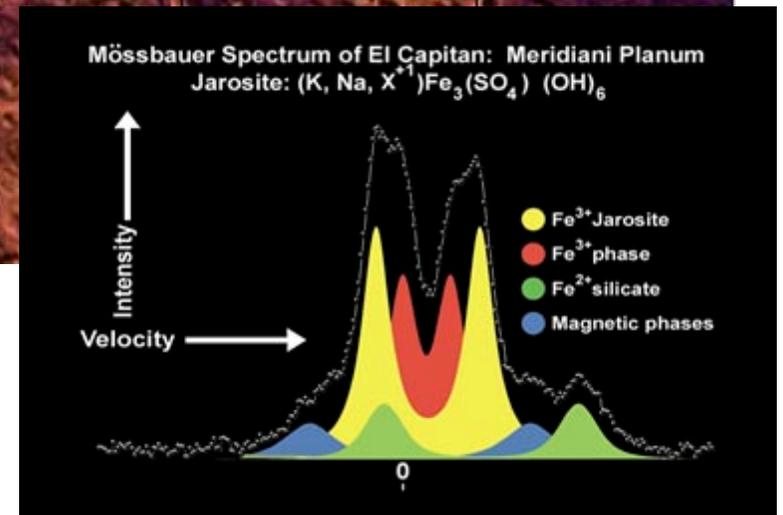
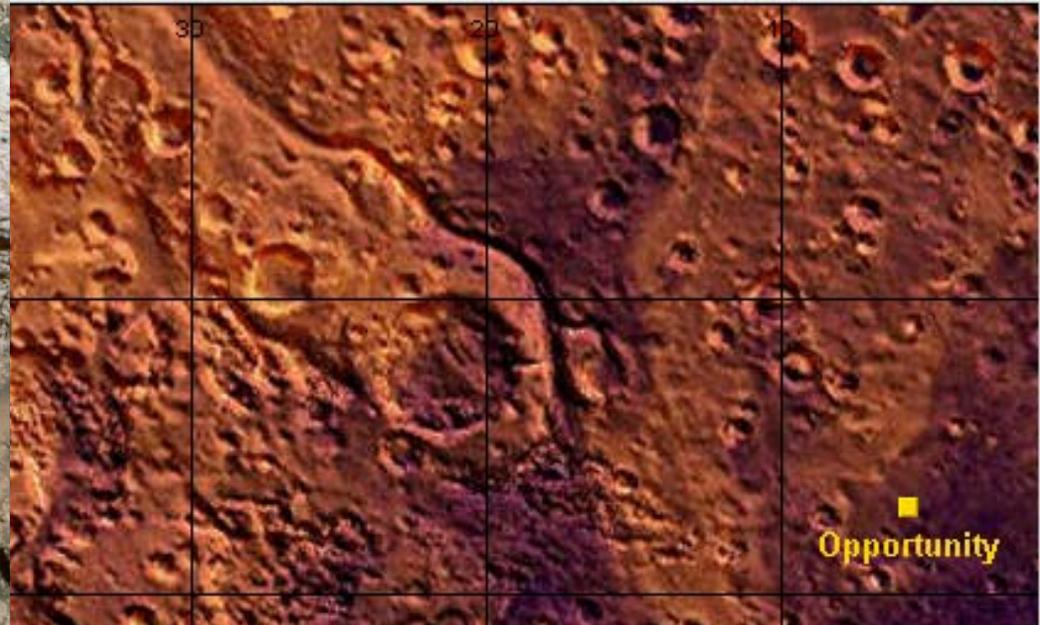
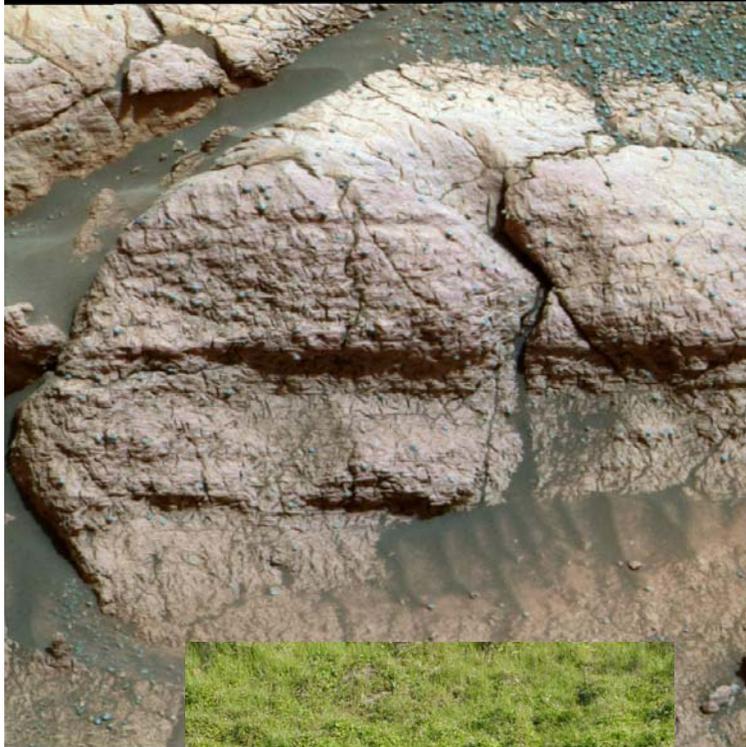
- Fluvial network with its origin in Arabia Terra
- Fan facies materials at the end of the valleys
- Delta deposits
- Streamlined island
- Tiu Valles seems to cut Ares Valles

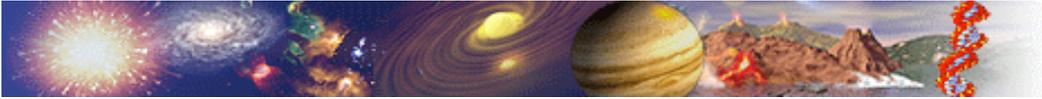
Geologic history

- Late Hesperian/Early Amazonian debris flows
- Ejecta from Chryse Basin became water-saturated
- Marsquakes liquefied the ejecta
- Flows into the basin

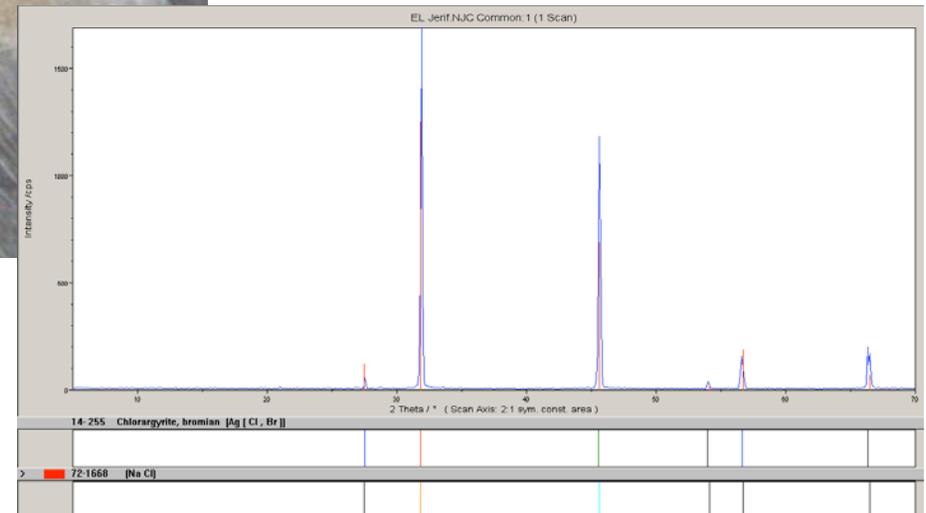
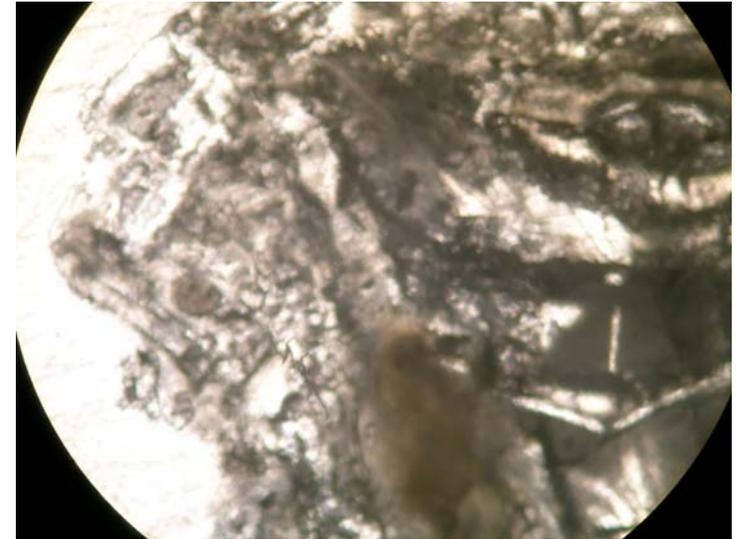
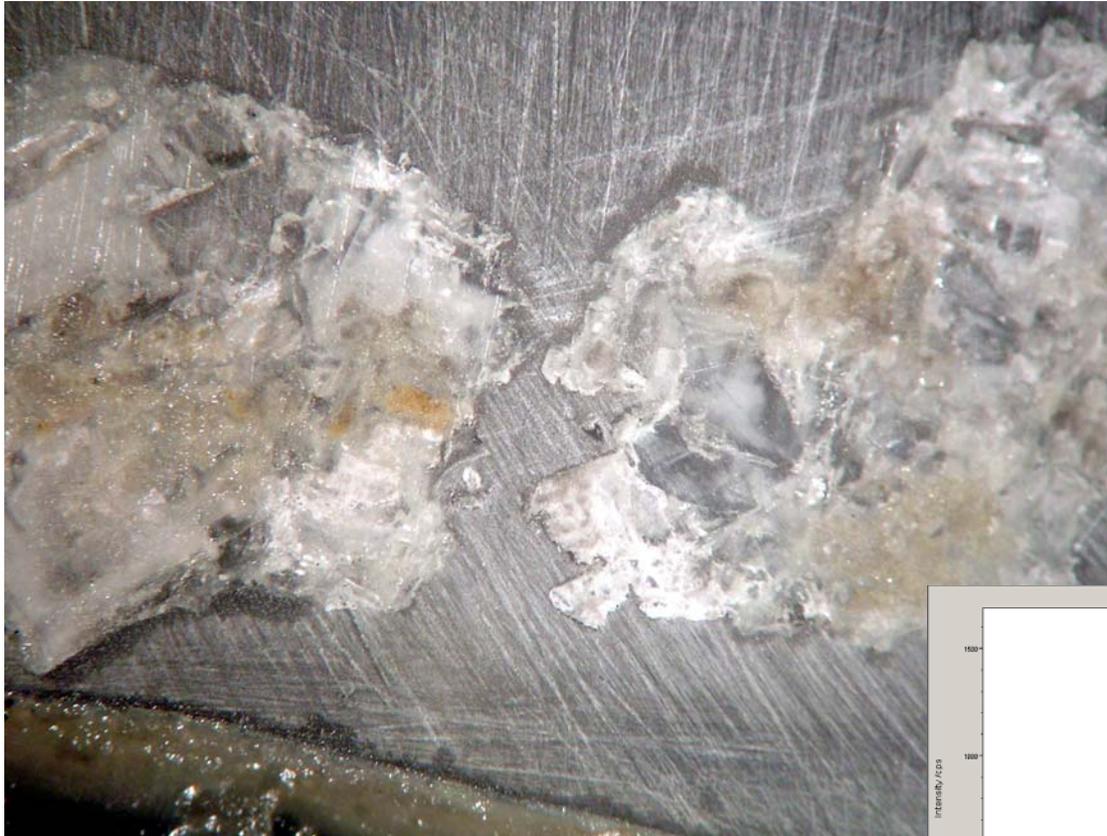


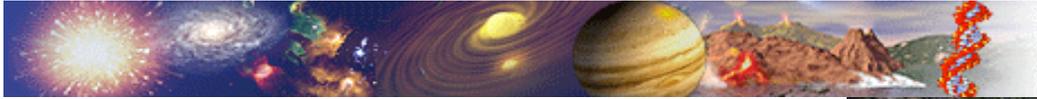
Mineralogical context



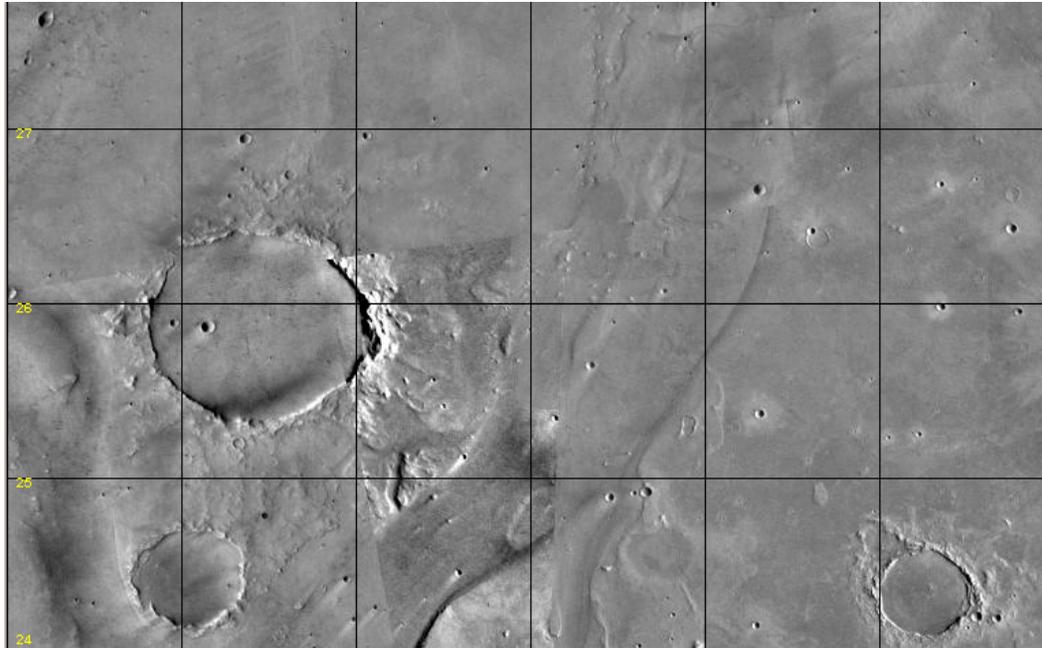
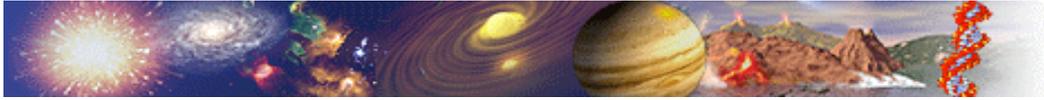


Organics inside salts





Mineralogical similarities: in this chemolithotrophic iron driven ecosystem, these minerals under anoxic conditions are the bioproducts and/or metabolites for habitability

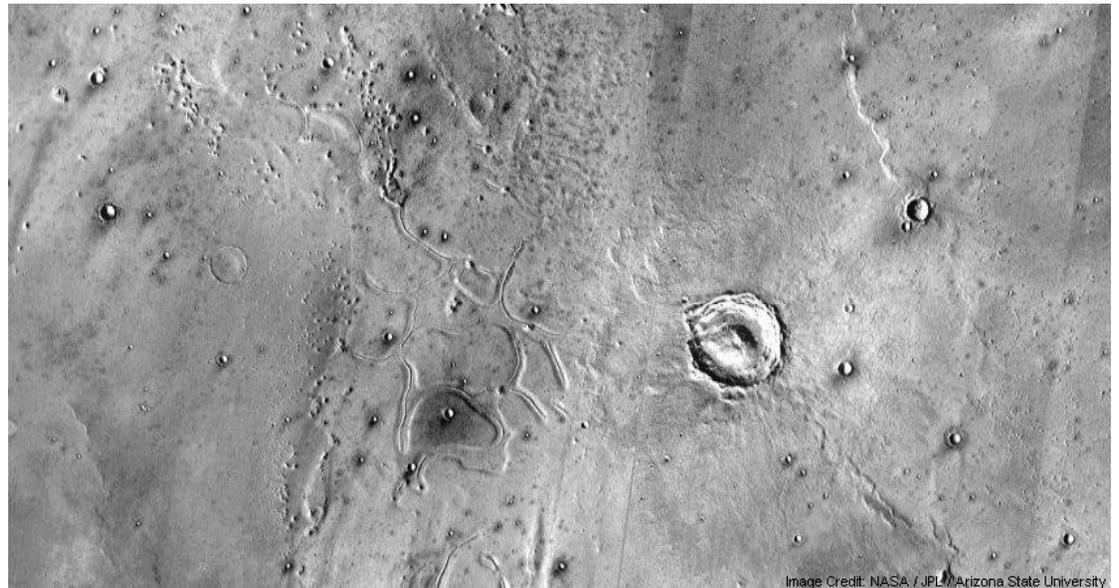
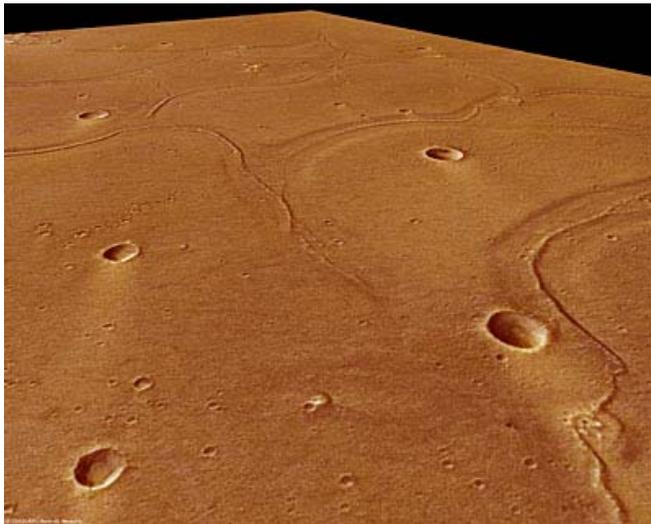


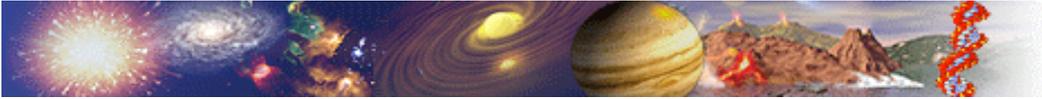
Sequence of fluvial processes

Sedimentary deposits

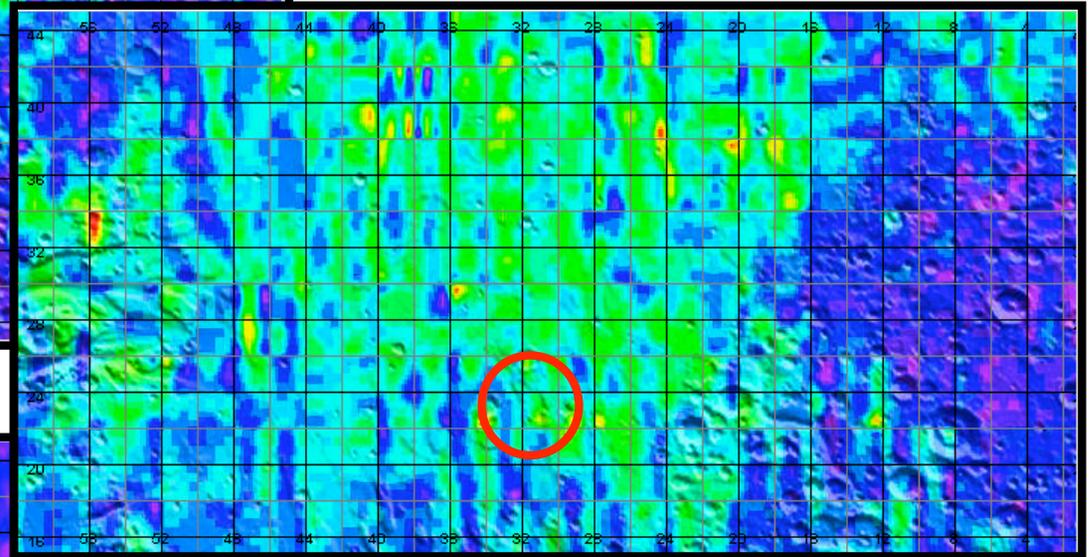
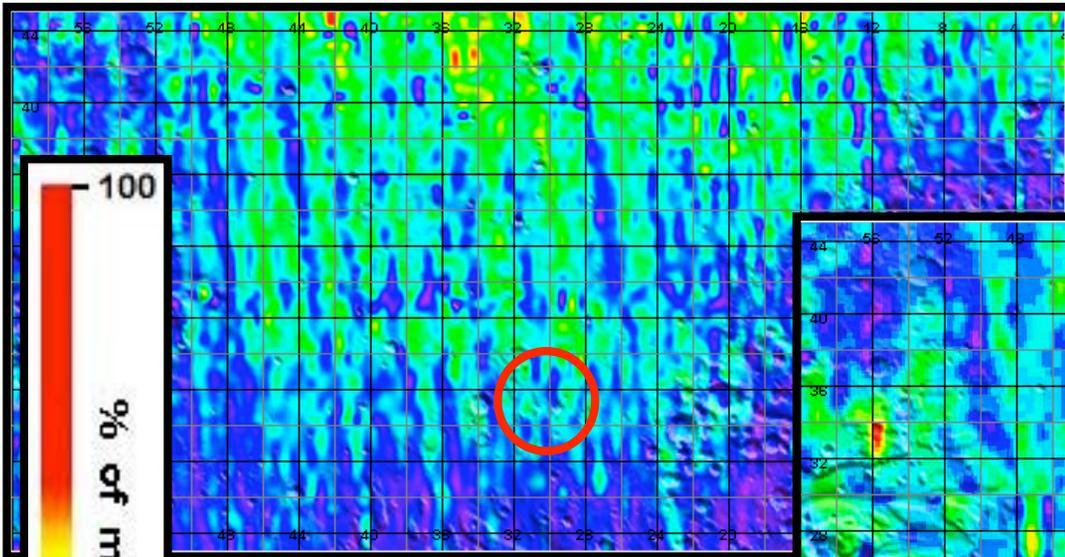
Depositional like structures

Long-term water

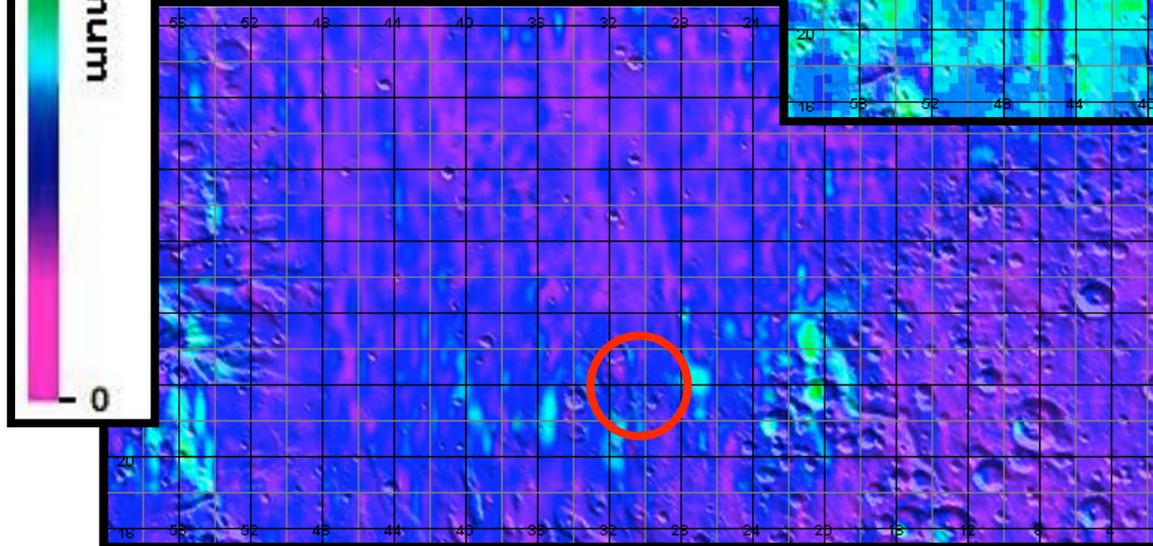




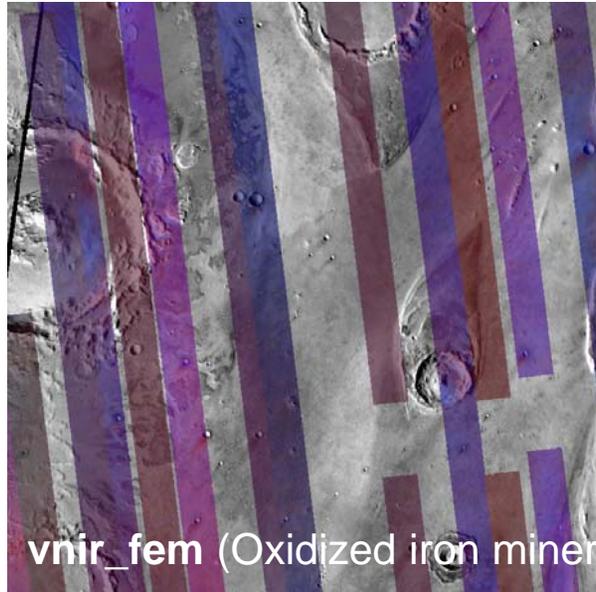
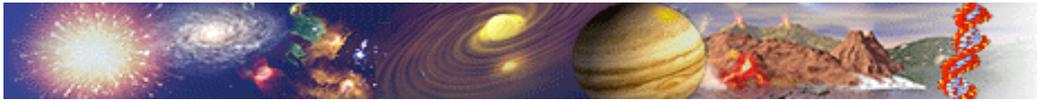
TES Andesite
(J.Banfield et al.)



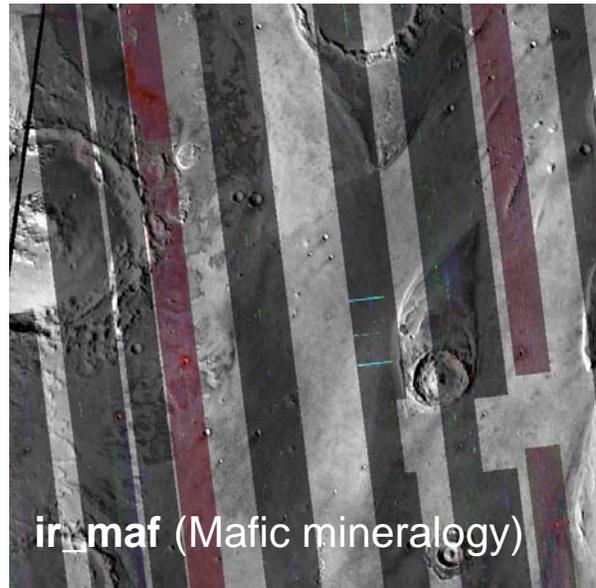
TES Hematite
(J.Banfield et al.)



TES Basalt
(J.Banfield et al.)

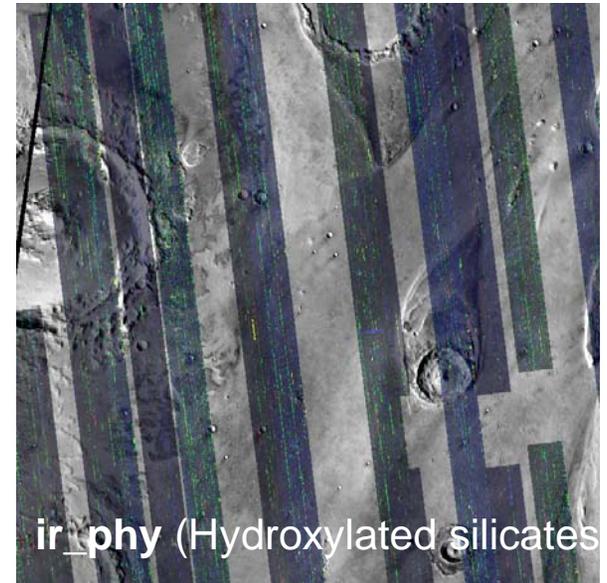


vni_fem (Oxidized iron minerals)

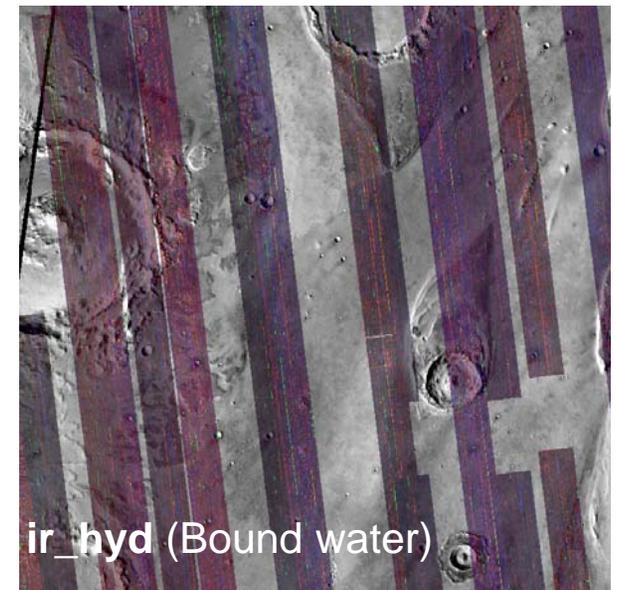


ir_maf (Mafic mineralogy)

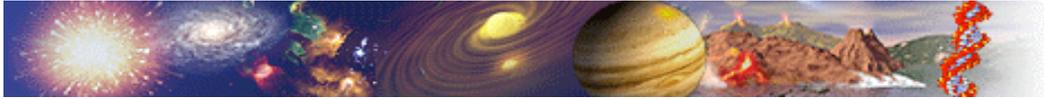
CRISM data at several spectral bands with surface composition: low percentage presence of olivine and pyroxene (ir_maf) but with a very interesting presence of altered materials in the form of water bound minerals like aluminum phyllosilicates or hydrated silica (ir_phy), and of very interest, the high presence of minerals or glasses with bound or dissolved molecular water and sulfates (ir_hyd) which represent altered materials and water



ir_phy (Hydroxylated silicates)



ir_hyd (Bound water)



Why Tiu Valles?

Preservation of biosignatures

Habitability

mineralogical similarities with
chemolithotrophic environments

could be Tiu Valles the subsurface
materials of Terra

Meridiani? Subsurface
protected environments

Surface radiation vs. Iron dust radiation
protection

past water presence

Mineralogical characterization and comparison

with MER Opportunity landing place
(Early Mars evolution?)

MSL objectives

Determine whether life ever arose on Mars

Biological potential

Past habitability: water and
mineralogical
evolution, Life
building blocks

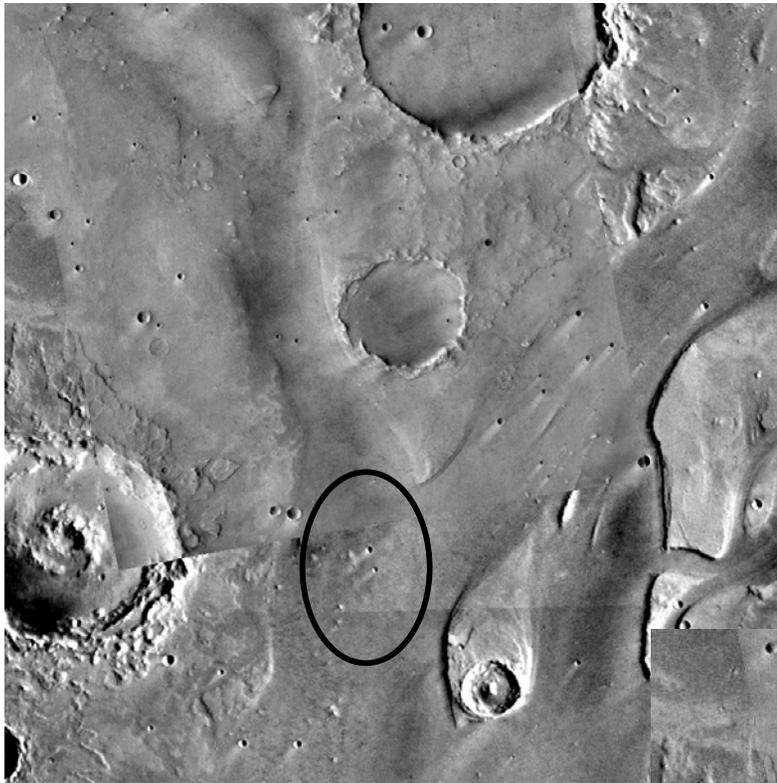
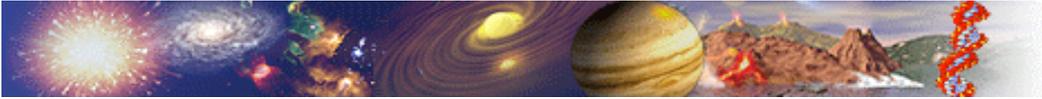
Characterize the climate of Mars

Humidity, T^a , P, UV, winds

Characterize the geology of Mars

Geology and geochemistry:
Organics?

Prepare for human exploration



Landing ellipse centered at 22.9N 32.25 W

MOLA elevation: -3.8 Km

Averaged Thermal inertia: $\sim 400 \text{ J/m}^2 \text{ K s}^{1/2}$

Expected temp at noon on landing season: 225 K

Averaged winds:

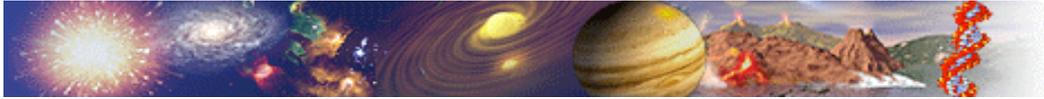
Northward @ 7 Km: $\sim 7 \text{ m/s}$

Eastward @ 7 Km: $\sim 2 \text{ m/s}$

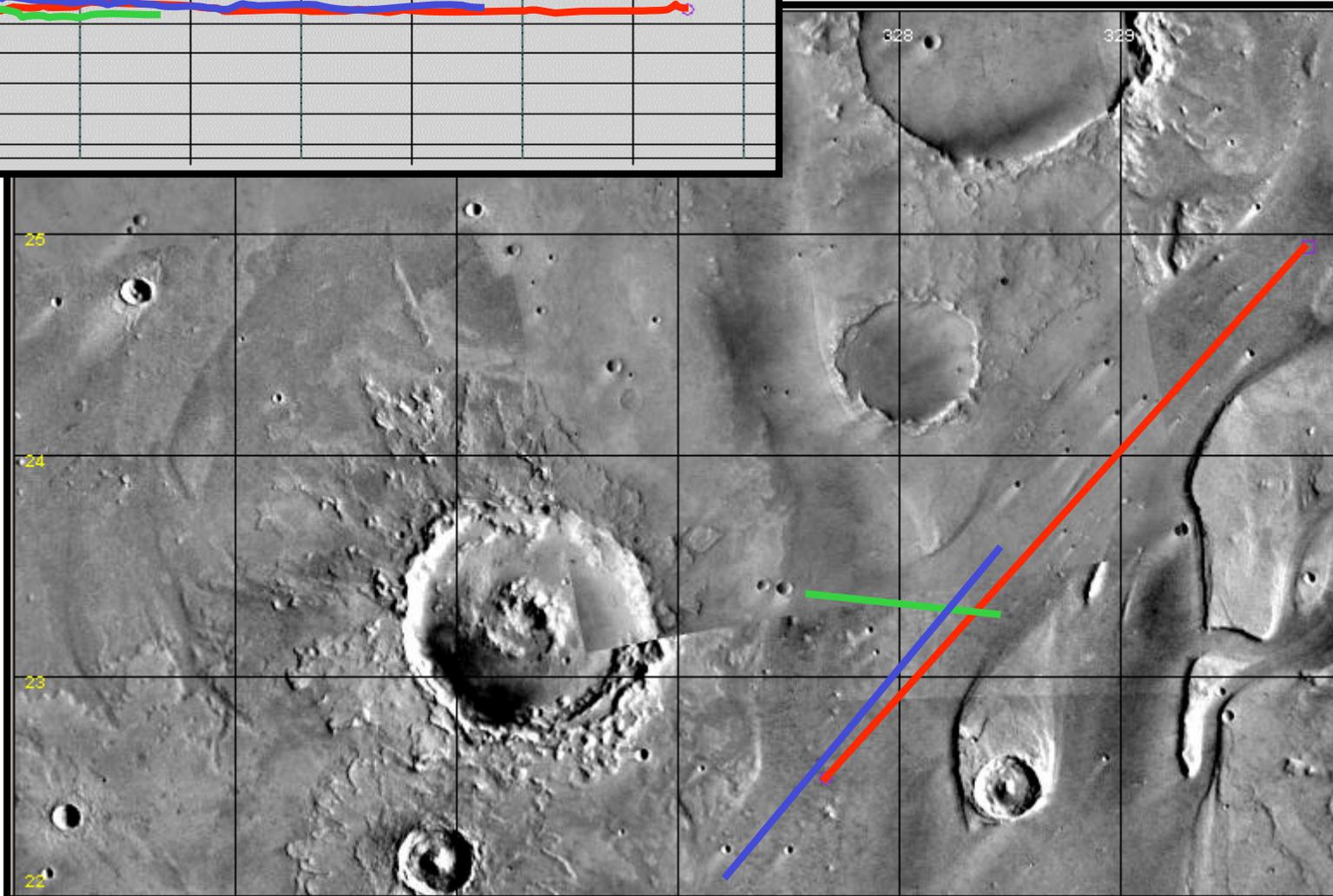
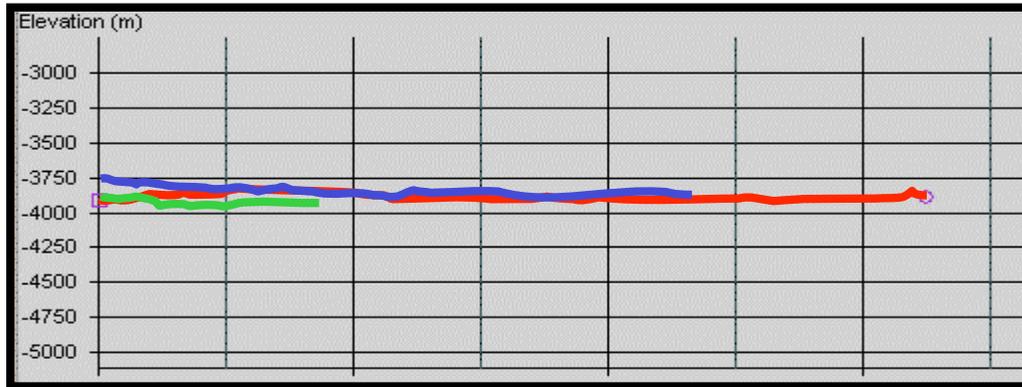
Northward @ 18 Km: $\sim 4 \text{ m/s}$

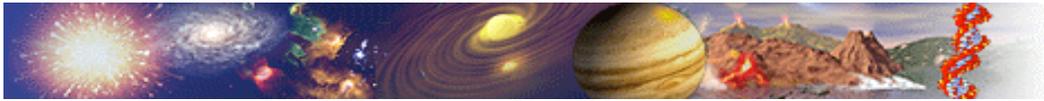
Eastward @ 18 Km: $\sim 4 \text{ m/s}$



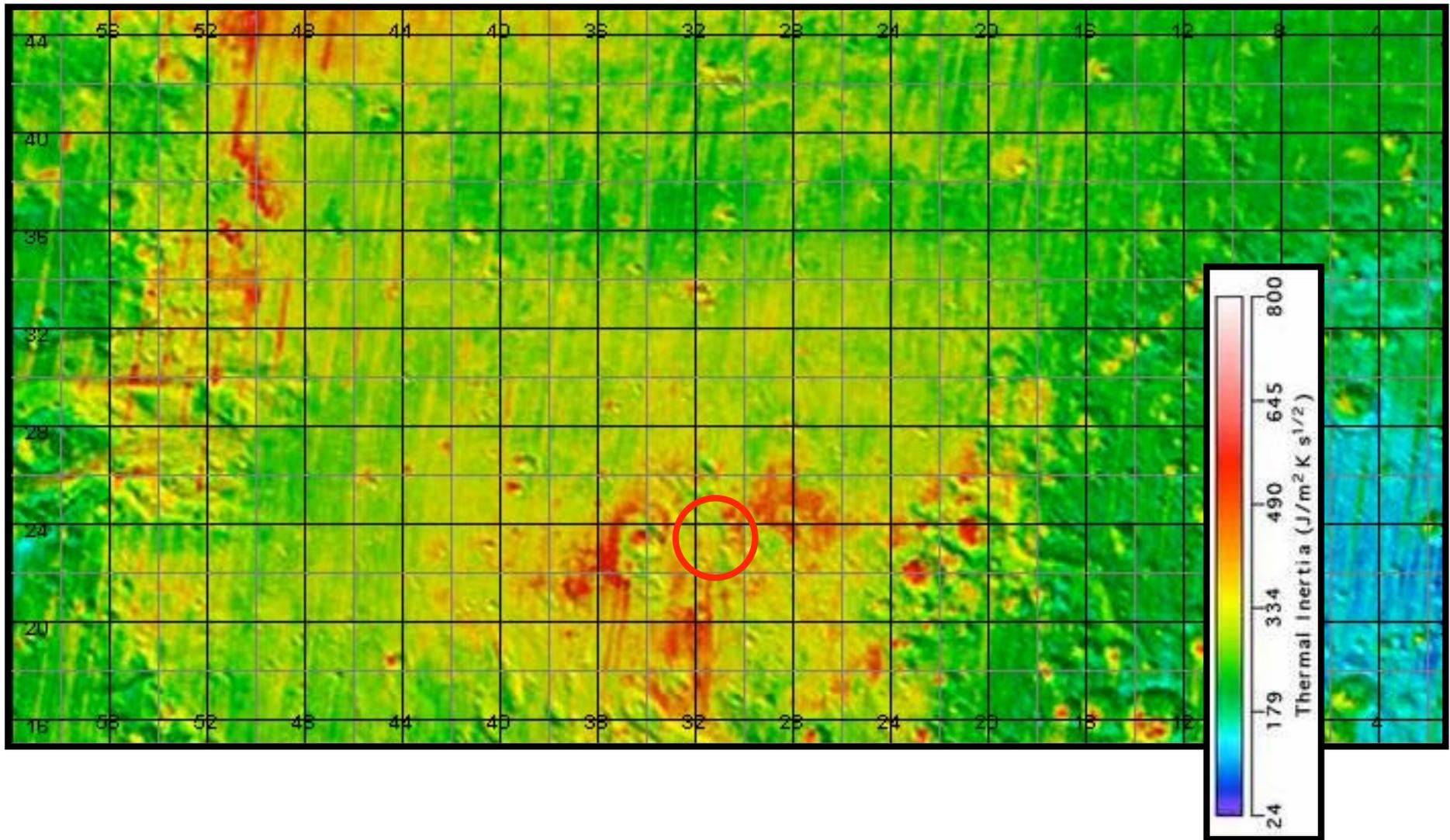


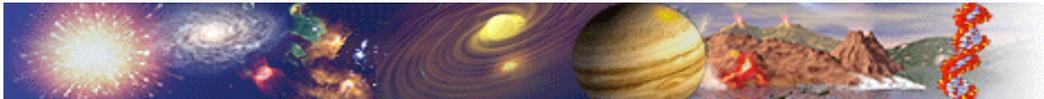
Elevation profiles from MOLA





Thermal inertia from TES





MOC SPO249401

Latitude: 24.62°

Longitude: 32.45° W

Resolution: 5.29 m

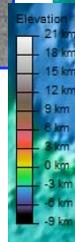
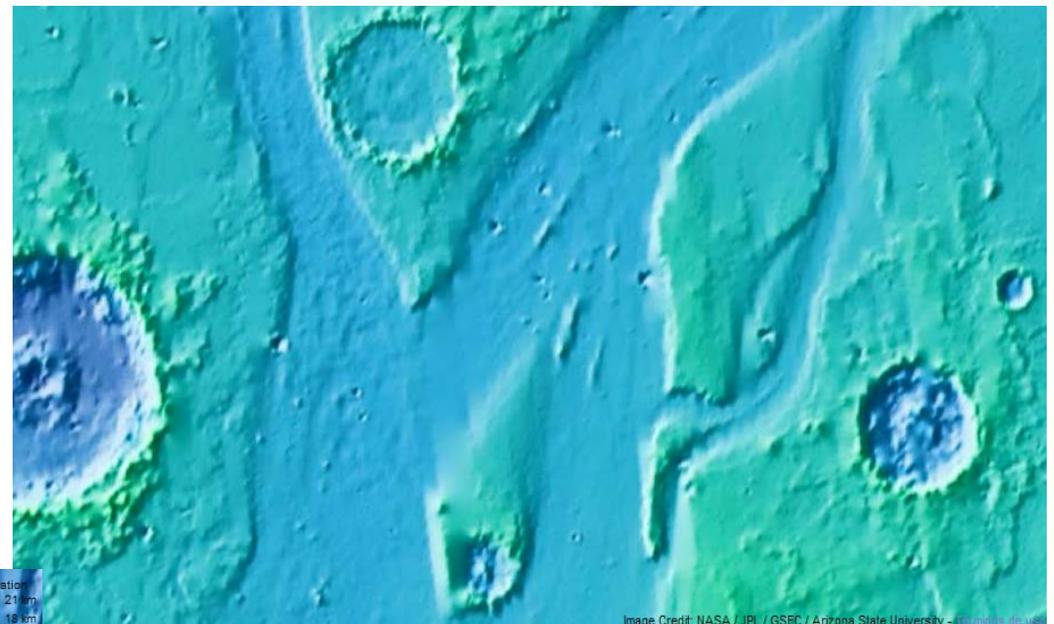
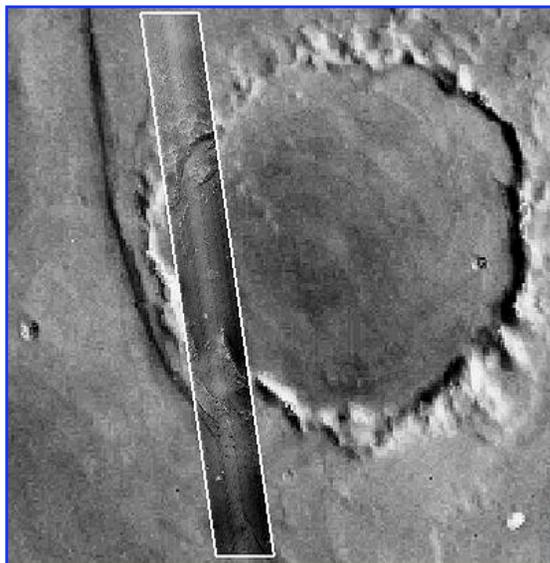
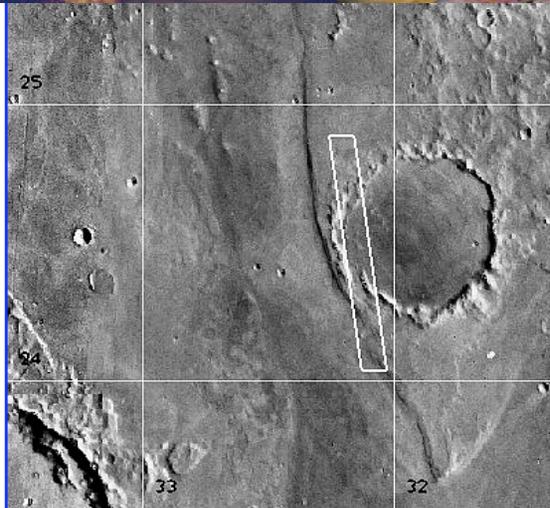
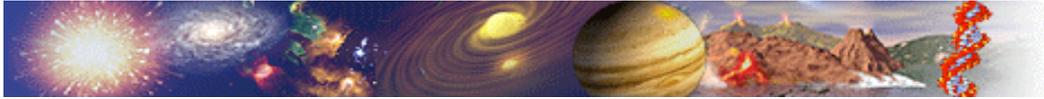


Image Credit: NASA / JPL / GSEC / Arizona State University - www.asu.edu



MSL payload and habitability study on Tiu Valles

Is there water on the subsurface?

DAN

Water signatures on rocks and surface soil?

MastCam

Is the presence of iron minerals ubiquitous on this area?

CheMin

Sediments origin? Nature of the deposits-textural information

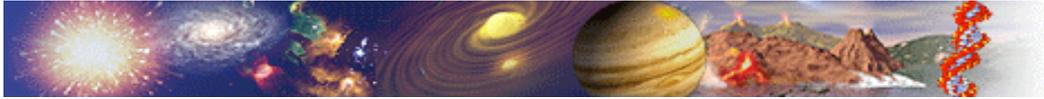
MastCam, ChemCam, CheMin

Are organics present?

SAM

Are environmental conditions suitable for habitability?

RFMS



Tiu Valles from the MSL objectives point of view

Biological objectives:

- Determine the nature and inventory of organic carbon compounds
- Inventory the chemical building blocks of life (carbon, hydrogen, nitrogen, oxygen, phosphorous and sulfur)
- Identify features that may represent the effects of biological processes

Geological and geochemical objectives:

- Investigate the chemical, isotopic, and mineralogical composition of the Martian surface and near-surface geological materials
- Interpret the processes that have formed and modified rocks and soils

Planetary process objectives:

- Assess long-timescale (i.e., 4-billion-year) atmospheric evolution processes
- Determine present state, distribution and cycling of water and carbon dioxide

Surface radiation objective:

- Characterize the broad spectrum of surface radiation, including galactic cosmic radiation, solar proton events and secondary neutrons